

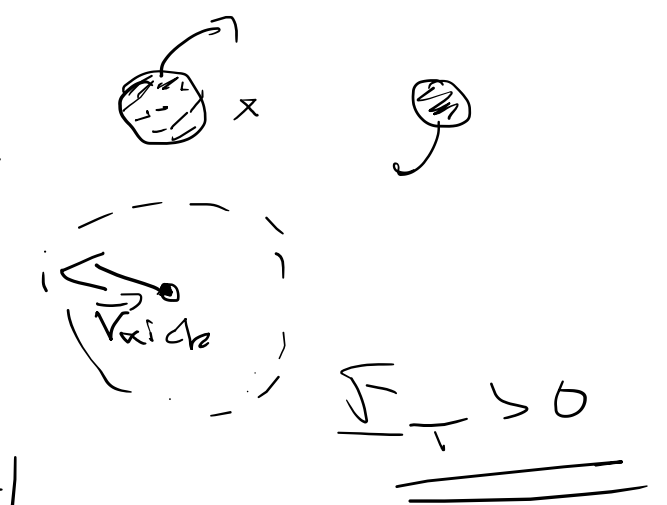


Модуль. Астрофизика.

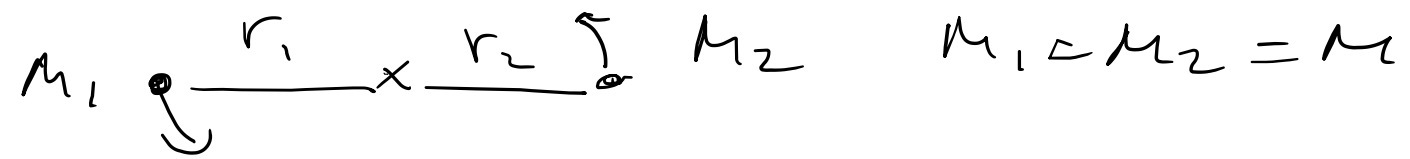
Семинар 5. (27.11)

① Pachung glow sw ~90%
 a) kick. Accumulation SW.

$v_{rock} \sim \text{few } 100 \text{ km/s}$



b). CD for $\omega \approx 10^4 \text{ km}$ $v_{rock} \ll v_{orb}$



$2a = r_{max} + r_{min} = l_{cr}$
 $a = 2r$

$$P^2 = \frac{4\pi^2 a^3}{G(M_1 + M_2)} = \frac{4\pi^2 8r^3}{2GM}$$

$$v_{orb} = \frac{2\pi r}{P}$$

$$v_{orb}^2 = \frac{1}{4} \frac{GM}{r}$$

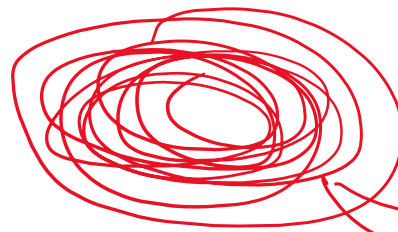
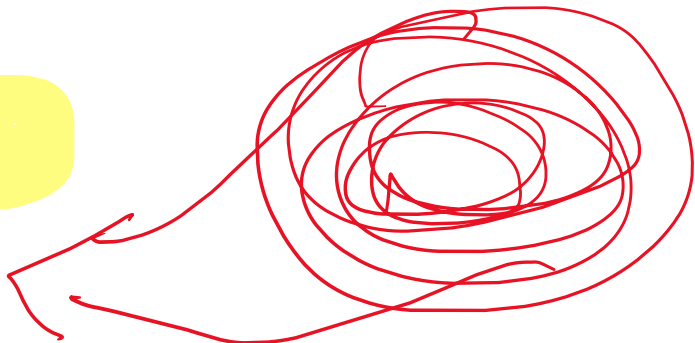
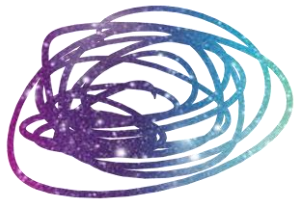
$$F_k = F_r^{(1)} + F_r^{(2)} = \frac{GM^2}{4r}$$

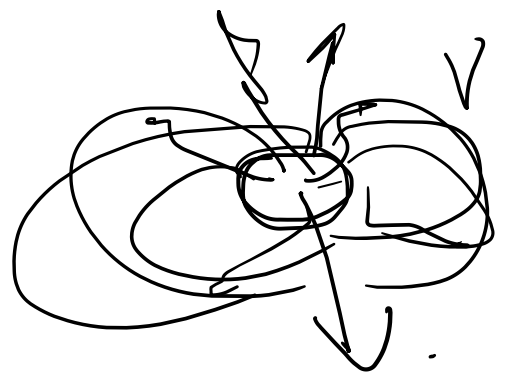
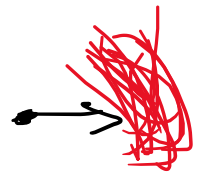
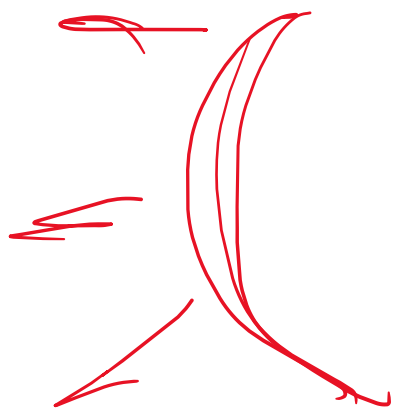
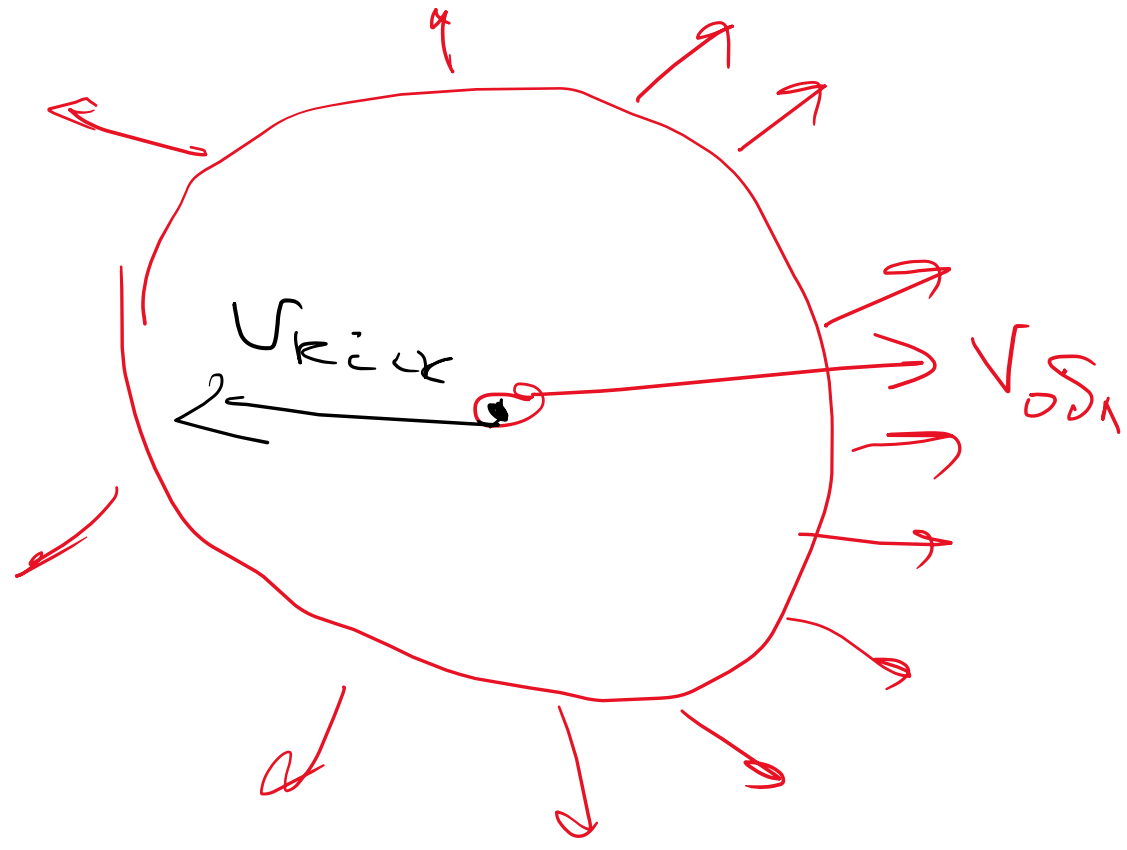
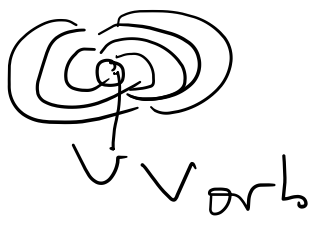
$$F_p = -\frac{GM^2}{2r}$$

$$F_T = F_k + F_p = -\frac{1}{4} \frac{GM^2}{r}$$

$\Rightarrow \Delta M = \frac{1}{2} M_{total}, \text{ so}$


$$F_T = 0$$



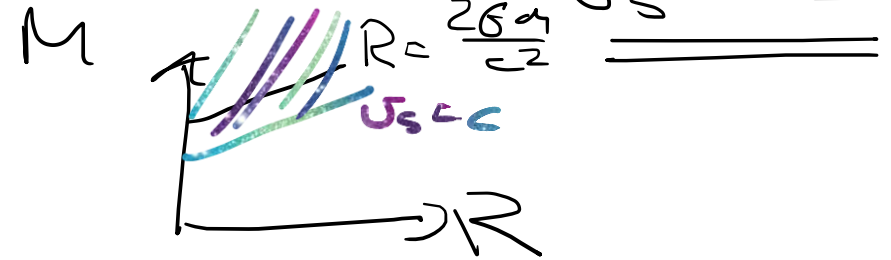


2. Фигурная масса галактик. 3,2 M_⊙
Ruffini



x  $M > 3 M_{\odot} \Rightarrow \text{BH}$

$\rho \uparrow$ $v_s \uparrow$
 $v_s = c$



$$a = a_1 + a_2$$

$$a_1 M_1 = a_2 M_2$$

$$a = a_1 \left(1 + \frac{a_2}{a_1} \right) = \frac{a_1}{M_2} (M_1 + M_2)$$

$$p^2 = \frac{4\pi^2 a^3}{G(M_1 + M_2)} = \frac{4\pi^2}{G(M_1 + M_2)} \frac{a_1^3}{M_2^3} (M_1 + M_2)^3$$

$$v_1 = \frac{2\bar{v} a_1}{P} \quad K = a_1 \cdot \frac{2\bar{v}}{P} \cdot \underline{\underline{\sin^2 i}}$$

$$a_1 = \frac{P}{2\bar{v}} \frac{1}{\sin^2 i} \quad K$$

$$P^2 = \frac{4\bar{v}^2}{G} \frac{(M_1 + M_2)^2}{M_2^3} \cdot \frac{P^3}{8\bar{v}^3} \frac{1}{\sin^3 i} K^3$$

$$\boxed{\frac{P}{2\bar{v}} \cdot K \cdot \frac{1}{G}}$$

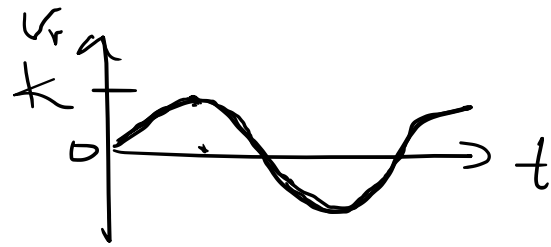
$$= \frac{M_2^3}{(M_1 + M_2)^2} \sin^3 i = f$$

$$q = \frac{M_1}{M_2}$$

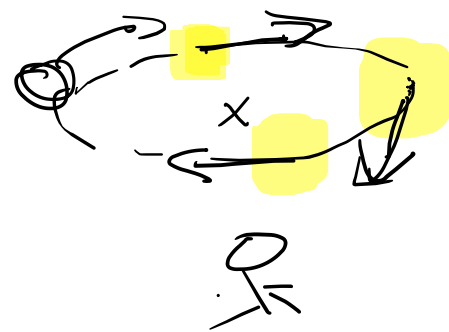
$$f = \frac{M_2}{(1+q)^3} \sin^3 i$$

$$M_2 = f \underbrace{(1+q)^3}_{> 1} \underbrace{\sin^3 i}_{> 1}$$

$$f = M_2 \frac{(1+q)^3}{\sin^3 i} \quad \text{OK}$$



K - nongamman.
nyetelhas
choposok



3. Интеграл МЗ (параметры, max. loss, асимптотика)

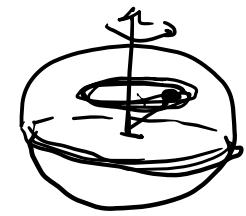
а) Вращение

$$\omega = 2\pi/P \quad P_{max}$$

$$\bar{F}_{Lmax} = \frac{I \omega^2}{2}$$



$$v_{ef} = \frac{2\pi R}{P_{max}} = \sqrt{\frac{GM}{R}}$$



$$I \sim MR^2$$

$$P_{max} = 2\pi \sqrt{\frac{R^3}{GM}}$$

$$P_{max} \approx 0.001^s$$

$$\frac{M}{R^3} \sim P$$

$$\frac{1}{P} = 716 T_1 \leftarrow \text{bad}$$

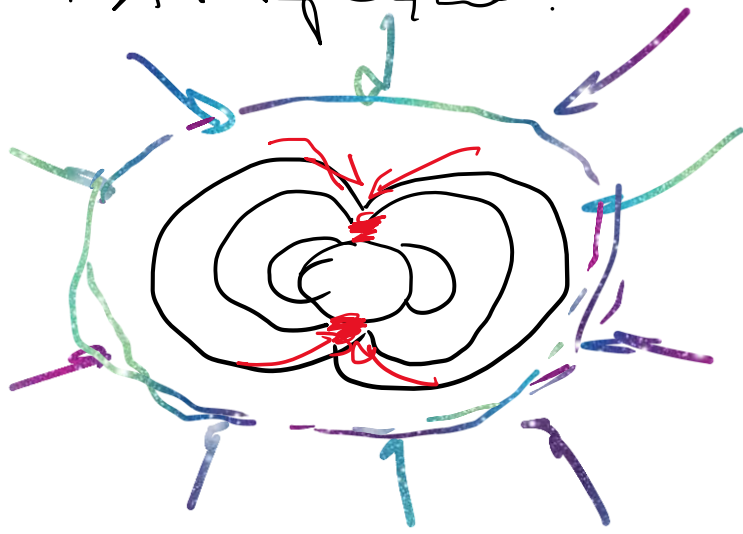
$$P_{max} = \frac{1}{\sqrt{g}}$$

$$I_{max} \sim 10^{45} \text{ g} \cdot \text{cm}^2$$

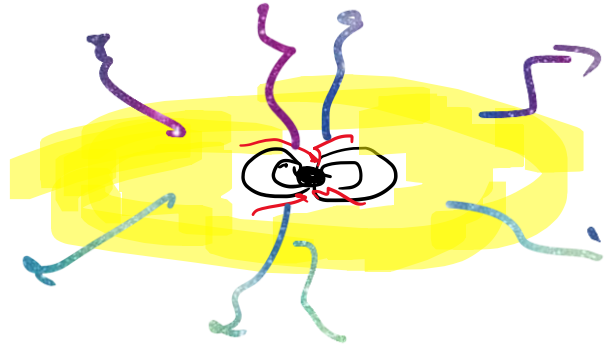
$$\bar{F}_{Lmax} = \frac{I \omega^2}{2} \frac{1}{P_{max}^2}$$

$$= 2 \cdot 10^{45} \cdot 10 \cdot \frac{1}{10^{-6}} = 2 \cdot 10^{52} \text{ dyn}$$

1. Anisotropy



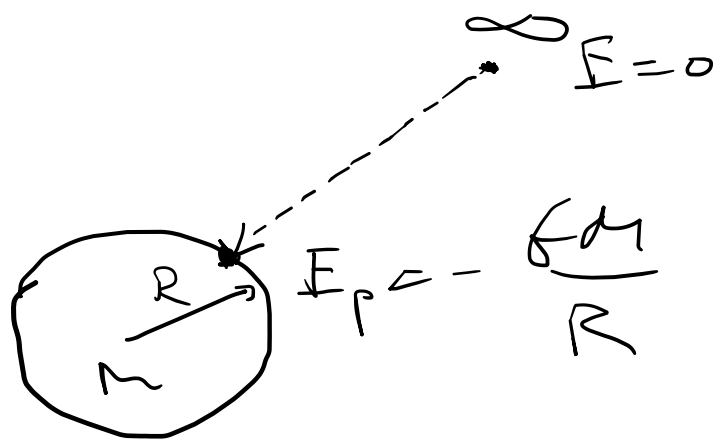
$$m \sim 1.9 \cdot 10^{20} \text{ g}$$



$$I_2 \rightarrow F \sim \frac{GM}{R}$$

$$\frac{10^{-7} \cdot 1.5 \cdot 2 \cdot 10^{33}}{1.5 \cdot 10^6} \approx 2 \cdot 10^{20} \text{ g}$$

$$\text{app. HJ} \approx \frac{10^0}{10^2} m c^2$$



2. Max. work

$$E = \frac{B^2}{8\pi}$$

$$B \sim \frac{1}{r^3}$$

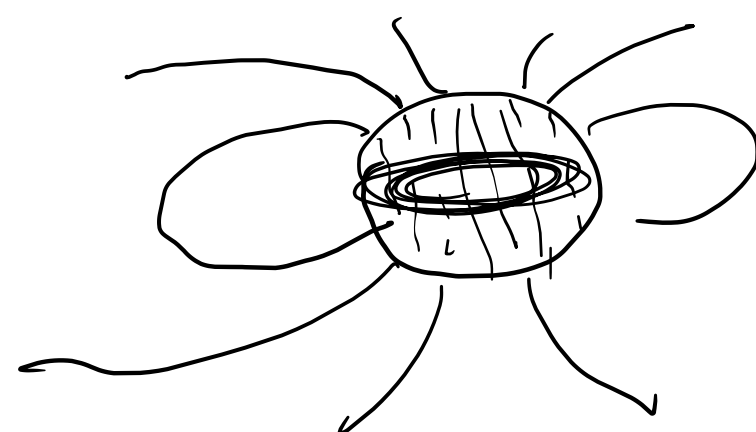
$$B_{\text{max}} = ?$$

$$\frac{B^2 R^3}{6} = \frac{GM^2}{R}$$

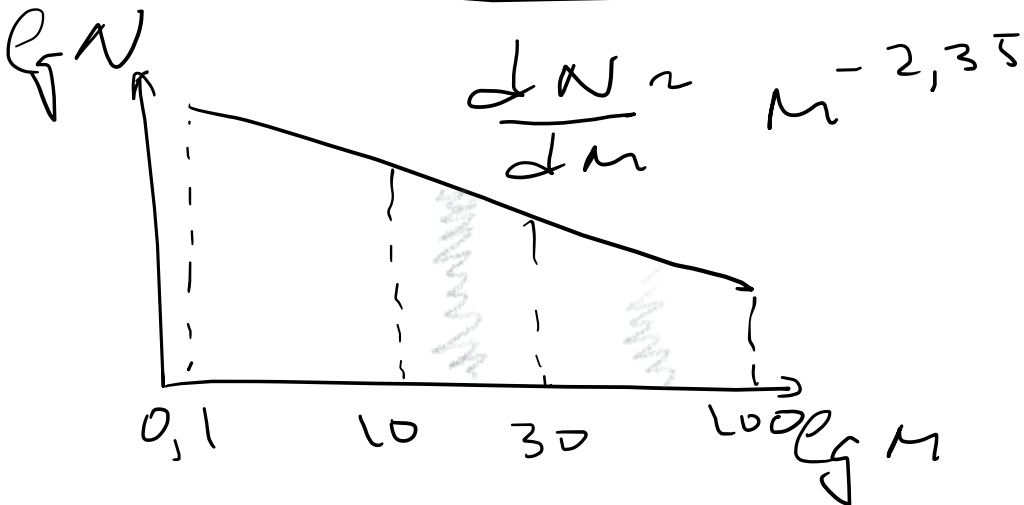
$$[B] = Tc$$

$$\rightarrow E_{\text{mag}} = \frac{B^2}{8\pi} \frac{4}{3} \pi R^3 = \frac{B^2 R^3}{6} = \frac{(10^{15} Tc)^2 (10^6 \text{ cm})^3}{6} \approx 1.5 \cdot 10^{47} \text{ erg}$$

$$B_{\text{max}} = \left[\frac{6GM^2}{R^2} \right]^{1/3} = \text{few} \cdot 10^{18} Tc$$

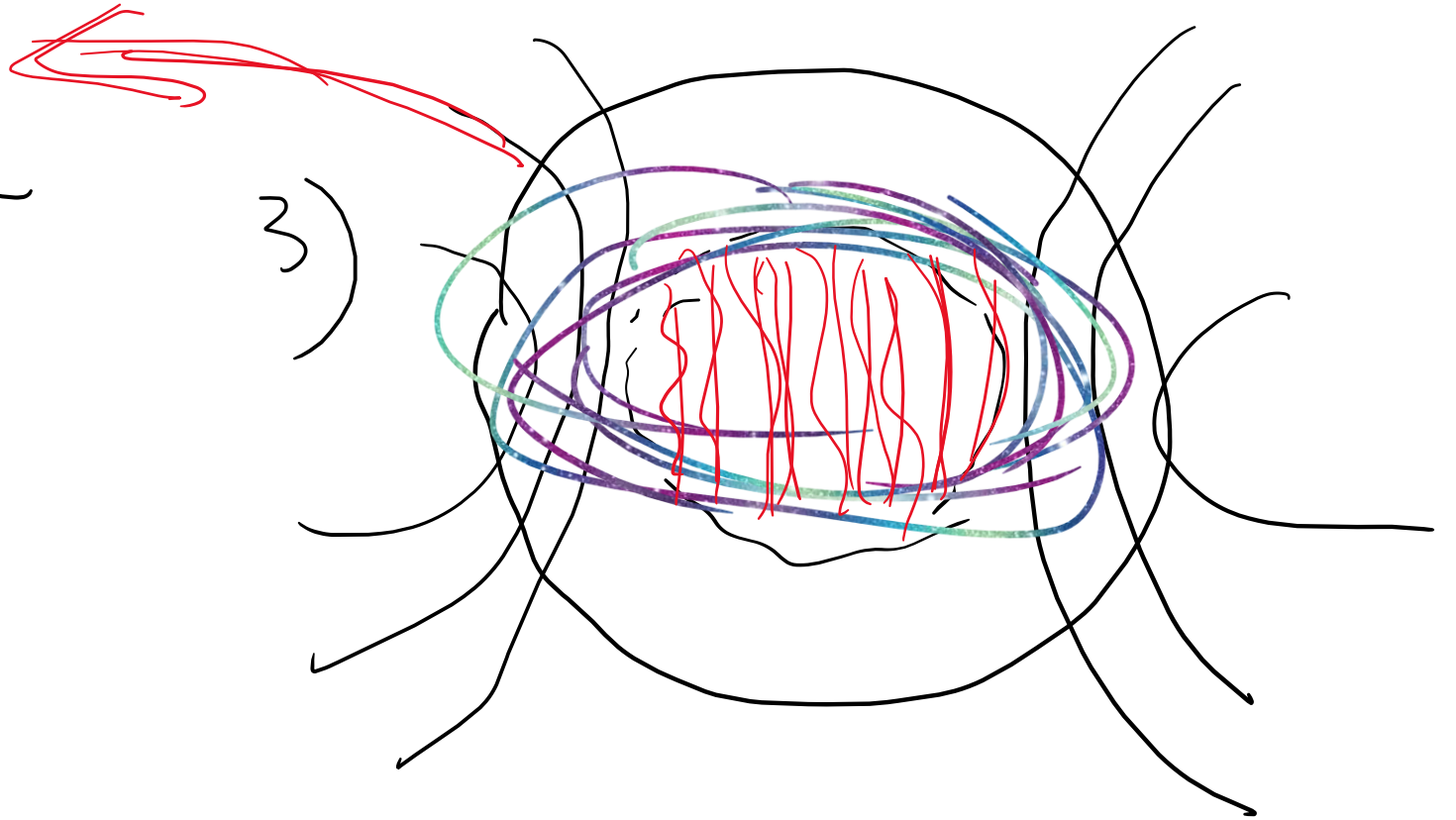
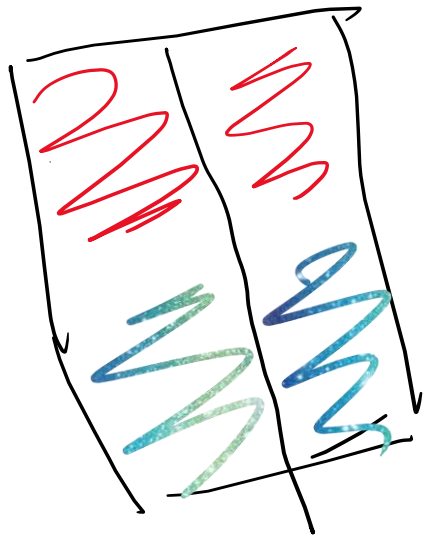
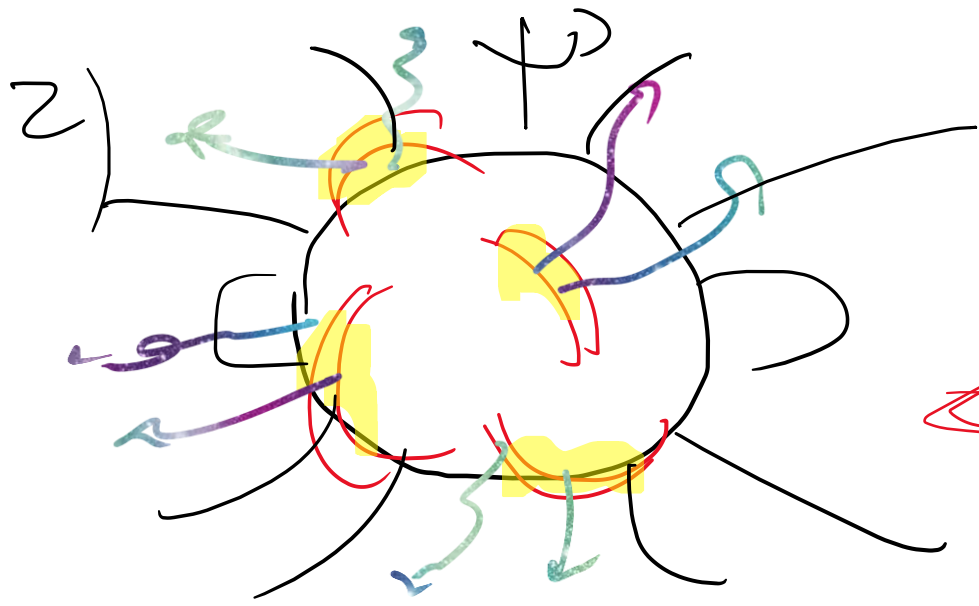


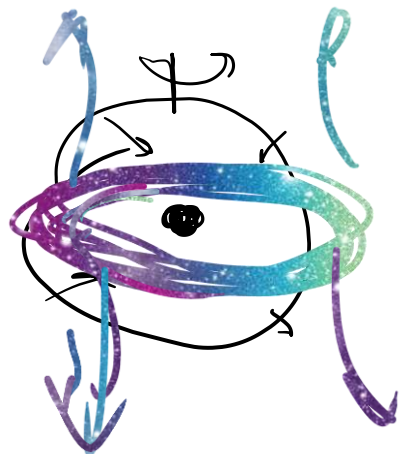
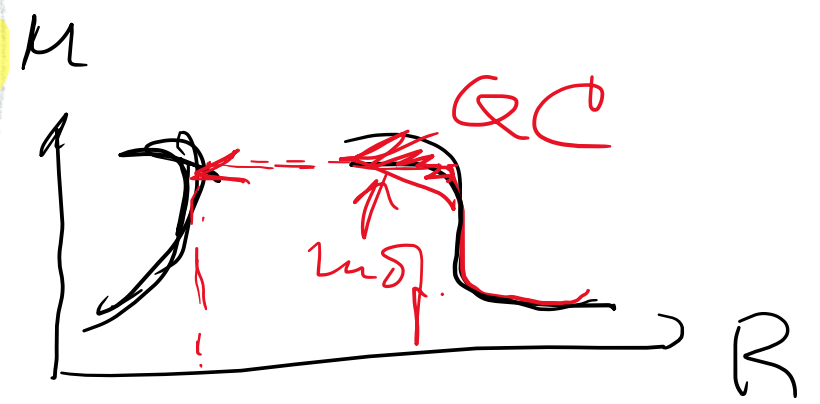
4. Donek NS - BK.

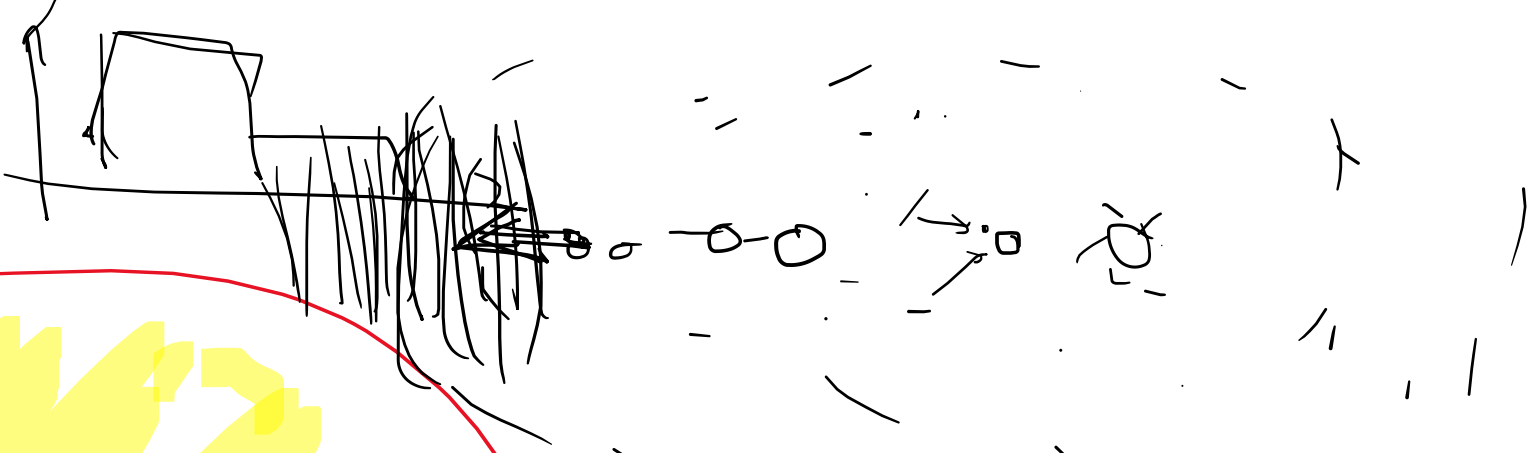
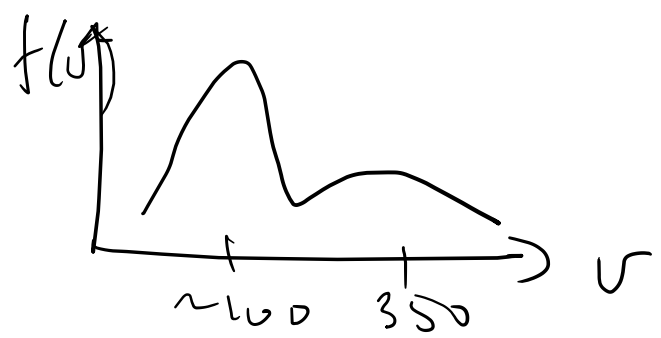


$$\langle a \rangle = \frac{A_1 \int + A_2 \int}{A_1 \int + A_2 \int}$$

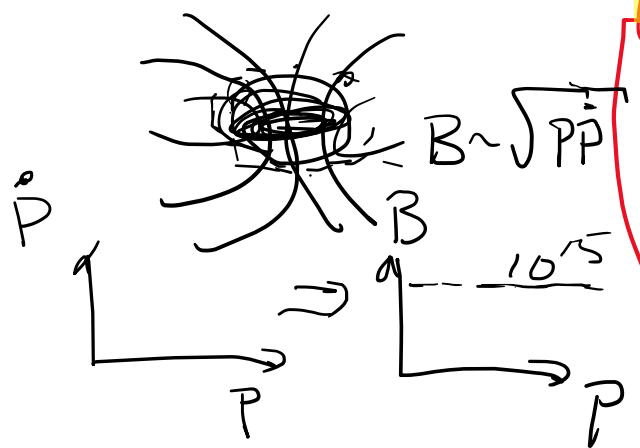
$$\frac{N_{NS}}{N_{BK}} = \frac{A \int_{10}^{30} M^{-2,35} dM}{A \int_{30}^{100} M^{-2,35} dM} \approx \left(\frac{30}{10} \right)^{1,35} \approx 4,5$$





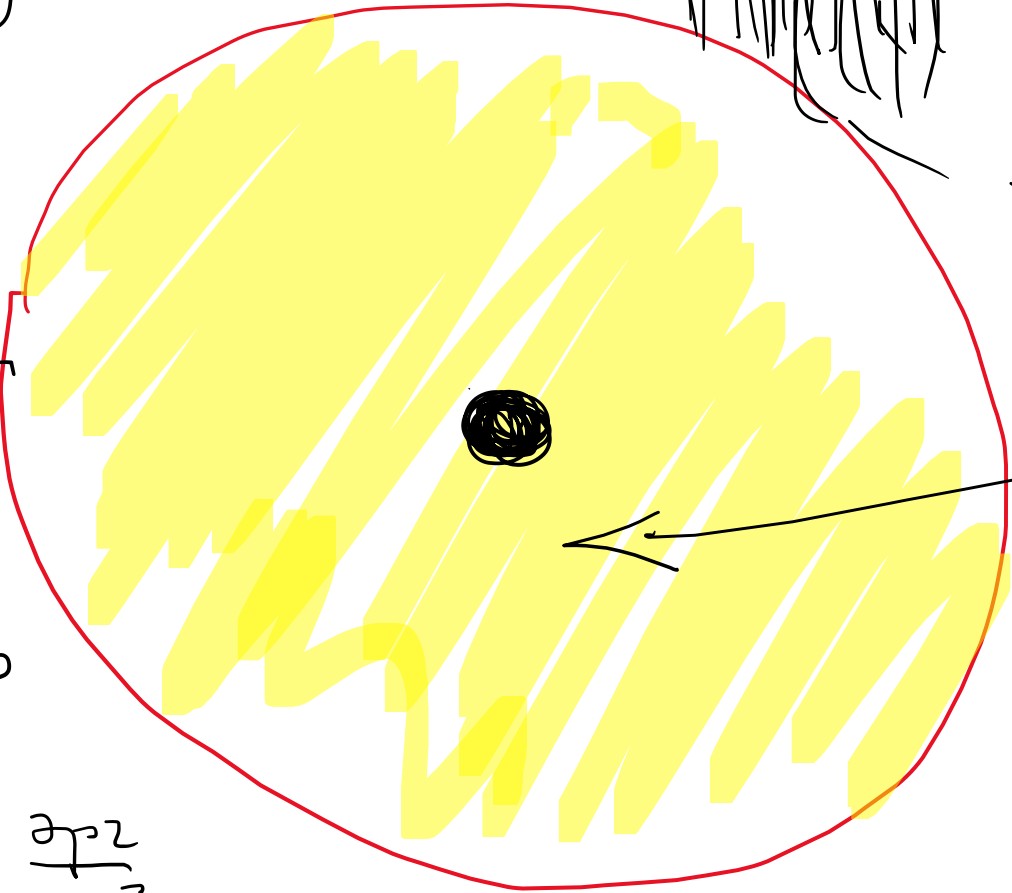


$$E_{\text{beam}} \sim B^2$$

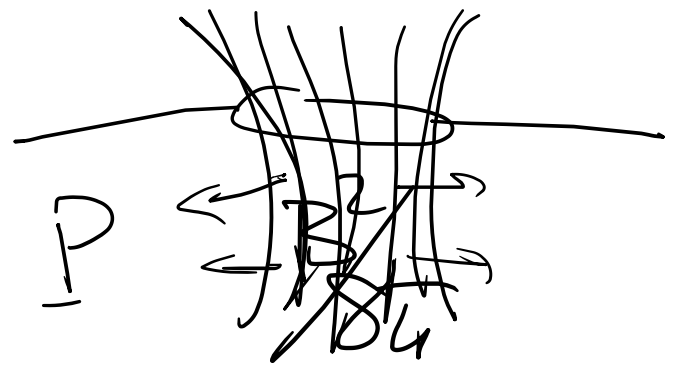


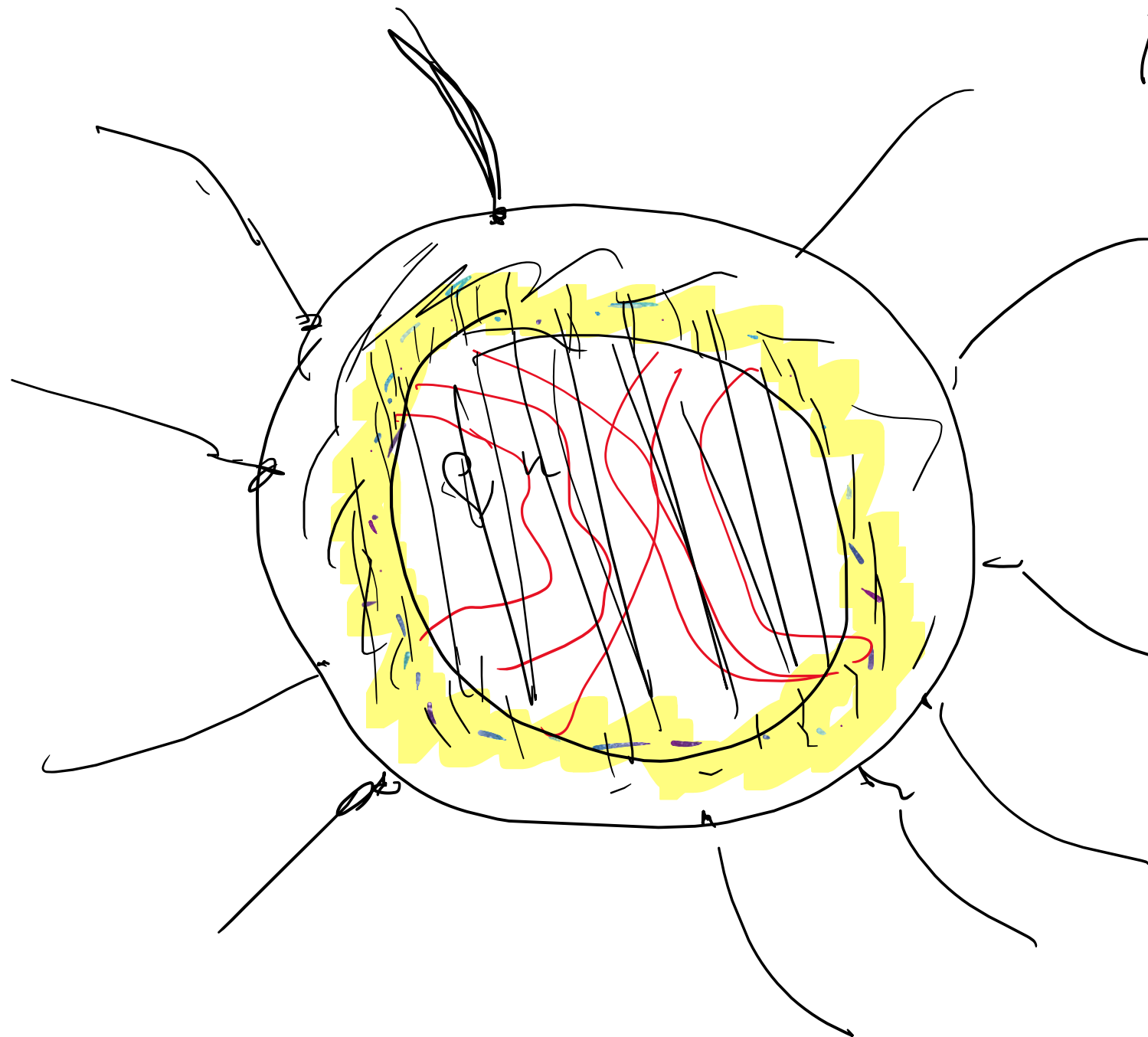
$$B \sim \sqrt{PP}$$

$$E = \frac{B^2}{8\epsilon_0} \rightarrow \frac{\text{J}}{\text{cm}^3}$$



$$\frac{B^2}{8\epsilon_0 c^2} \rightarrow \frac{2}{\text{cm}^3}$$





$$\frac{I_{sn}}{I} \sim 1,4\%$$

~~9~~

$$-\frac{\Delta V}{V} \rightarrow$$

$$V = \frac{1}{P}$$