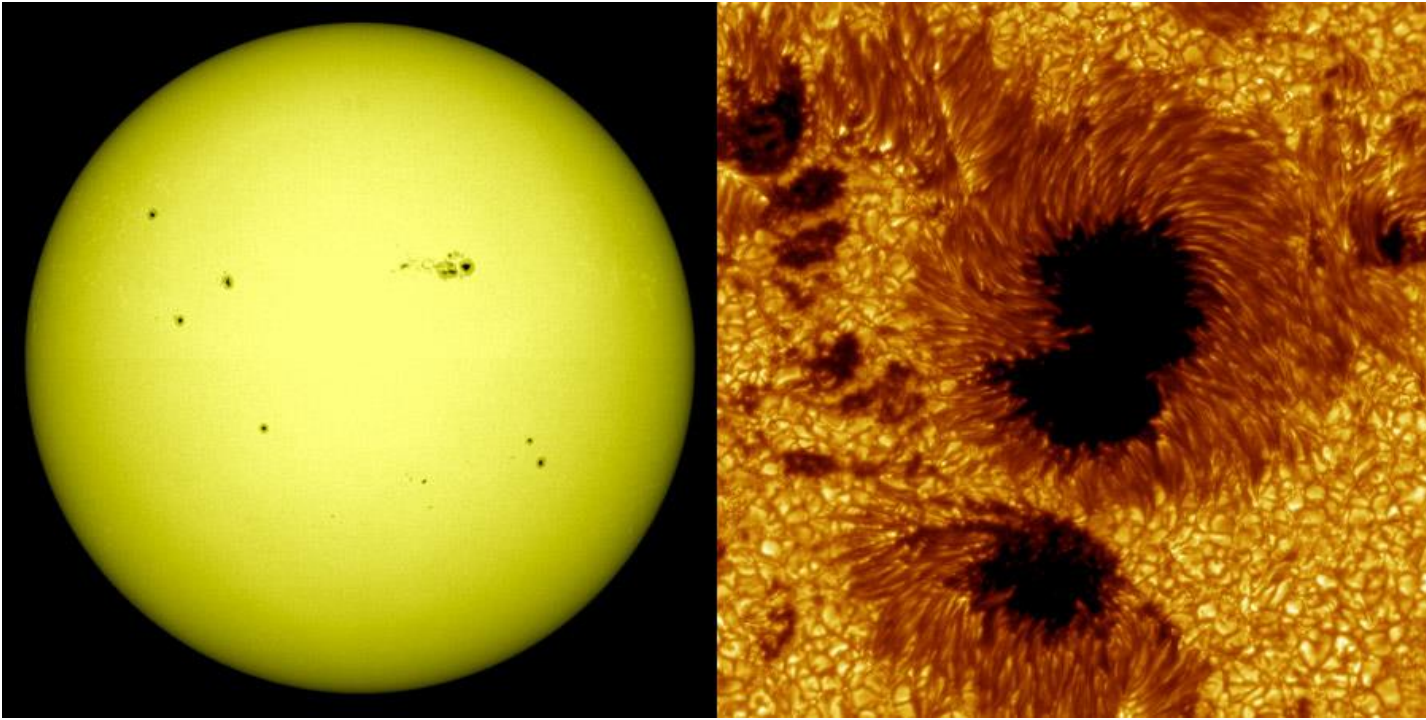




# Астрофізика и космологія.

Семинар 3. (30.01)



$$\bar{T} = 1500\text{K}$$

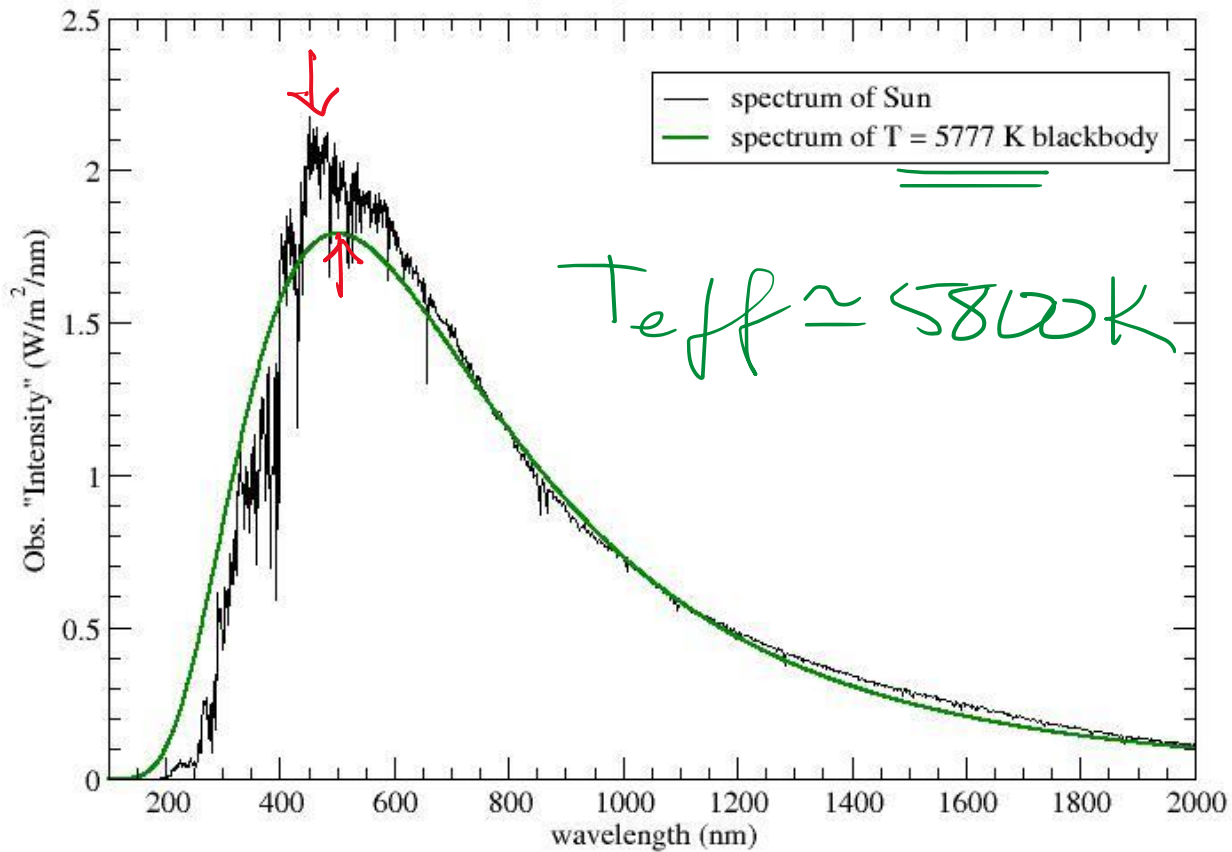
$$T_{\text{eff}} = 5800\text{K}$$

$$\frac{\langle T_{\text{in}} \rangle}{\langle T_{\text{eff}} \rangle} \sim \frac{T_{\text{in}}^4}{T_{\text{eff}}^4} \approx 0.3$$

$$\int B_{\nu} d\nu \sim T^4 \int x^3 e^{-x} dx \sim T^4$$

# Sun's Spectrum vs. Thermal Radiator

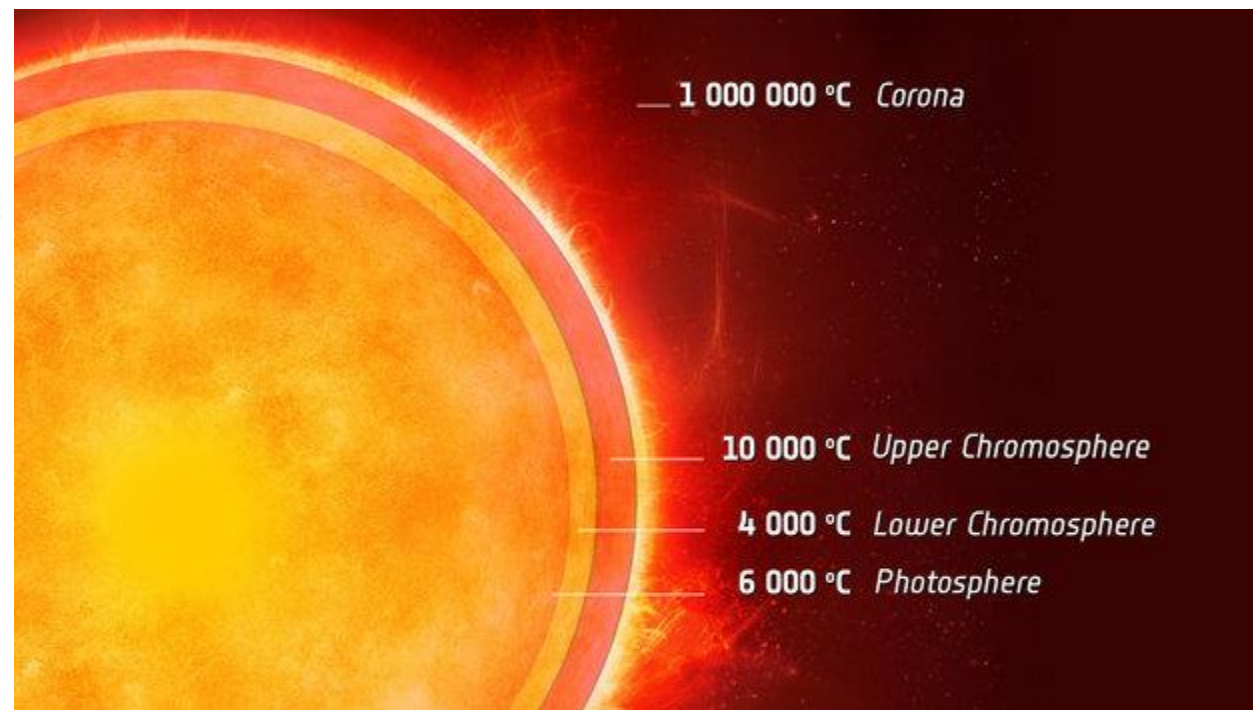
of a single temperature  $T = 5777\text{ K}$



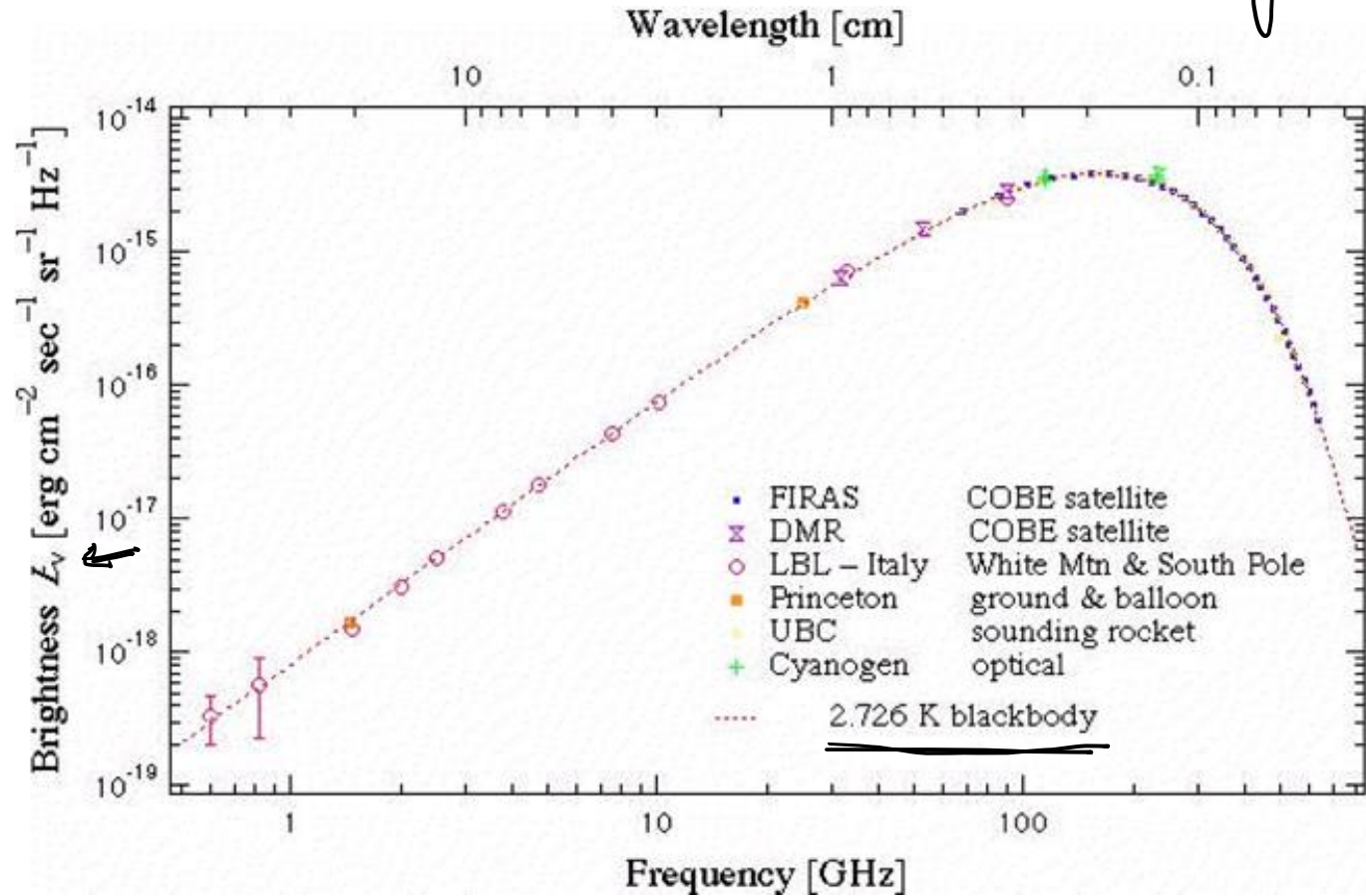
$$\lambda_{\text{max}} = 4600\text{ \AA}$$

$$\lambda_{\text{max}} = \frac{0,29(\text{cm})}{T}$$

$$T = \frac{0,29}{4,6 \cdot 10^{-5}} \approx 6300\text{ K}$$



Permanently uniform  
 Cosmic microwave background



$$T = 2,73 \text{ K}$$

$$B_\nu = \frac{2h\nu^3}{c^2} e^{-\frac{h\nu}{kT}}$$

$$\rightarrow B_\lambda = \frac{2hc^2}{\lambda^5} e^{-\frac{hc}{\lambda kT}}$$

$$\frac{dB_\nu}{dT} = 0$$

$$\frac{dB_\lambda}{d\lambda} = 0$$

$$\lambda_{max} = \frac{0,29 \text{ cm}}{T}$$

$$\lambda_{max} = \frac{0,29}{2,73} \approx 0,1 \text{ cm}$$

}  $B_\lambda$

$$E_r = \frac{c}{\pi^2 h^3} \int u p^3 d^3 p \quad p = \frac{h\nu}{c} \quad T \approx 3K$$

$$E_r = \frac{c(2\pi)^3}{\pi^2 h^3} \int u \left(\frac{h\nu}{c}\right)^3 \frac{h}{c} d\nu = \frac{8\pi h}{c^3} \int h\nu^3 d\nu$$

$$h = \frac{1}{e^{h\nu/kT} - 1} \quad N = \frac{E_r}{h\nu} = \frac{8\pi h}{c^3} \int \frac{\nu^3 d\nu}{h\nu [e^{h\nu/kT} - 1]} =$$

$$= \frac{8\pi}{c^3} \int_0^\infty \frac{\nu^2 d\nu}{e^{h\nu/kT} - 1} = \frac{8\pi}{c^3} \frac{(kT)^3}{h^3} \int_0^\infty \frac{x^2 dx}{e^x - 1}$$

$$\int_0^\infty \frac{x^2 dx}{e^x - 1} = \int_0^\infty \frac{x^2 dx}{e^x} = -x^2 e^{-x} \Big|_0^\infty - \int_0^\infty (-e^{-x}) 2x dx =$$

$$= 2 \int_0^\infty x e^{-x} dx = -2x e^{-x} \Big|_0^\infty - \int_0^\infty (-e^{-x}) 2 dx = 2 \int_0^\infty e^{-x} dx = 2$$

$$N = 2 \frac{8\pi}{c^3} \frac{(kT)^3}{h^3} = \frac{16\pi \cdot (1.38)^3 \cdot 27 \cdot 10^{-48}}{27 \cdot 10^{30} (6.63)^3 \cdot 10^{-81}} = \frac{16\pi (1.38)^3}{(6.63)^3} \cdot 1000 \approx 453$$

$$N \approx 400 \text{ cm}^{-3}$$

$$N \approx \frac{1}{\lambda_{\max}}$$

$$\lambda_{\max} = \frac{0.29 \text{ cm}}{T}$$

$$N \approx \frac{1}{2} \frac{1}{\lambda_{\max}}$$

$$N \approx T^{-3}$$

$\Phi_{\text{OH}}$

$n_{\lambda} - n_{\lambda} \frac{\partial n}{\partial \lambda}$   
 $\partial B \cdot \text{cm}^{-3}$

$n, \text{cm}^{-3}$

$P_{\text{aguo}}$

$10^{-8}$

$10^{-2}$

$\gamma \Phi_{\text{out}}$

$\sim 10^{-2}$

$\sim 10^{-2}$

$E_{ph}(\text{out}) \sim 1 \text{ dB}$

$P_{\text{entz}}$

$2 \cdot 10^{-5}$

$3 \cdot 10^{-9}$

$\chi$

$10^{-5}$

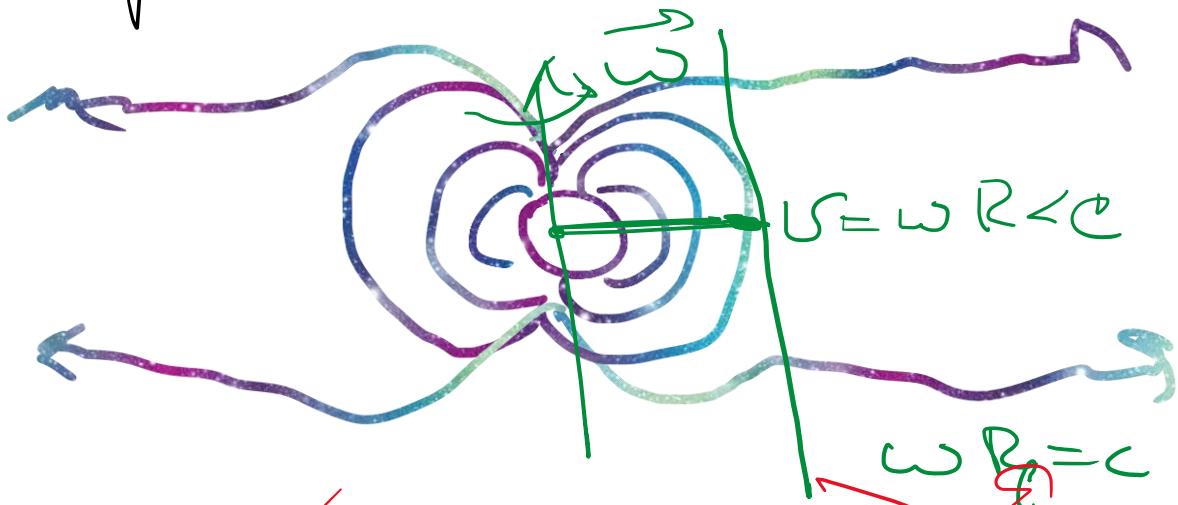
$\sim 10^{-11}$

$UK$

$4 \cdot 10^{-2}$

0,3

Эркт Терн.



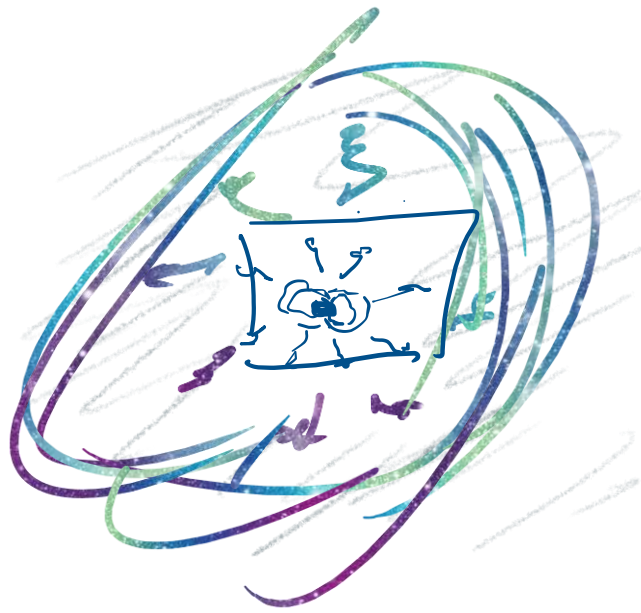
$R_{NS} = 10 \text{ km}$

$R_e = \frac{3 \cdot 10^{10} \cdot 0,633}{2\pi} = \frac{10^9}{2\pi} \approx 1,6 \cdot 10^8 \text{ cm}$

$\omega = \frac{2\pi}{P}$

$\frac{2\pi}{P} R_e = c$

$R_e = \frac{cP}{2\pi}$



ПЕРЛОТ

$$\underline{d} = 2200 \text{ nm}$$

$$F = 1 \int y = 10^{-23} \frac{\text{erg}}{\text{cm}^2 \cdot \text{c} \cdot \text{s}}$$

$$T_b = \frac{F}{\Omega} \cdot \frac{c^2}{2\nu^2 R}$$

$$T_b(1) = \frac{10^{-23} \cdot 10^{21} \cdot 4 \cdot 10^{43}}{2 \cdot \pi \cdot 10^{12} \cdot 10^{16} \cdot 1,58 \cdot 10^{16}} =$$

$$\approx \frac{1}{2} \frac{10^{41}}{10^{44}} \approx 5 \cdot 10^{28} \text{ K}$$

$$T_b(2) = 5 \cdot 10^{24} \text{ K}$$

$$I_{\text{nk}} = 206265 \text{ a.u.}$$

$$l_{\text{ae}} = 1,5 \cdot 10^{13} \text{ cm}$$

$$l_{\text{nk}} \approx 3,1 \cdot 10^{18} \text{ cm}$$

$$T_b = \frac{c^2}{2\nu^2 R} I_{\text{nk}}$$

$$\Omega = \frac{S}{r^2}$$

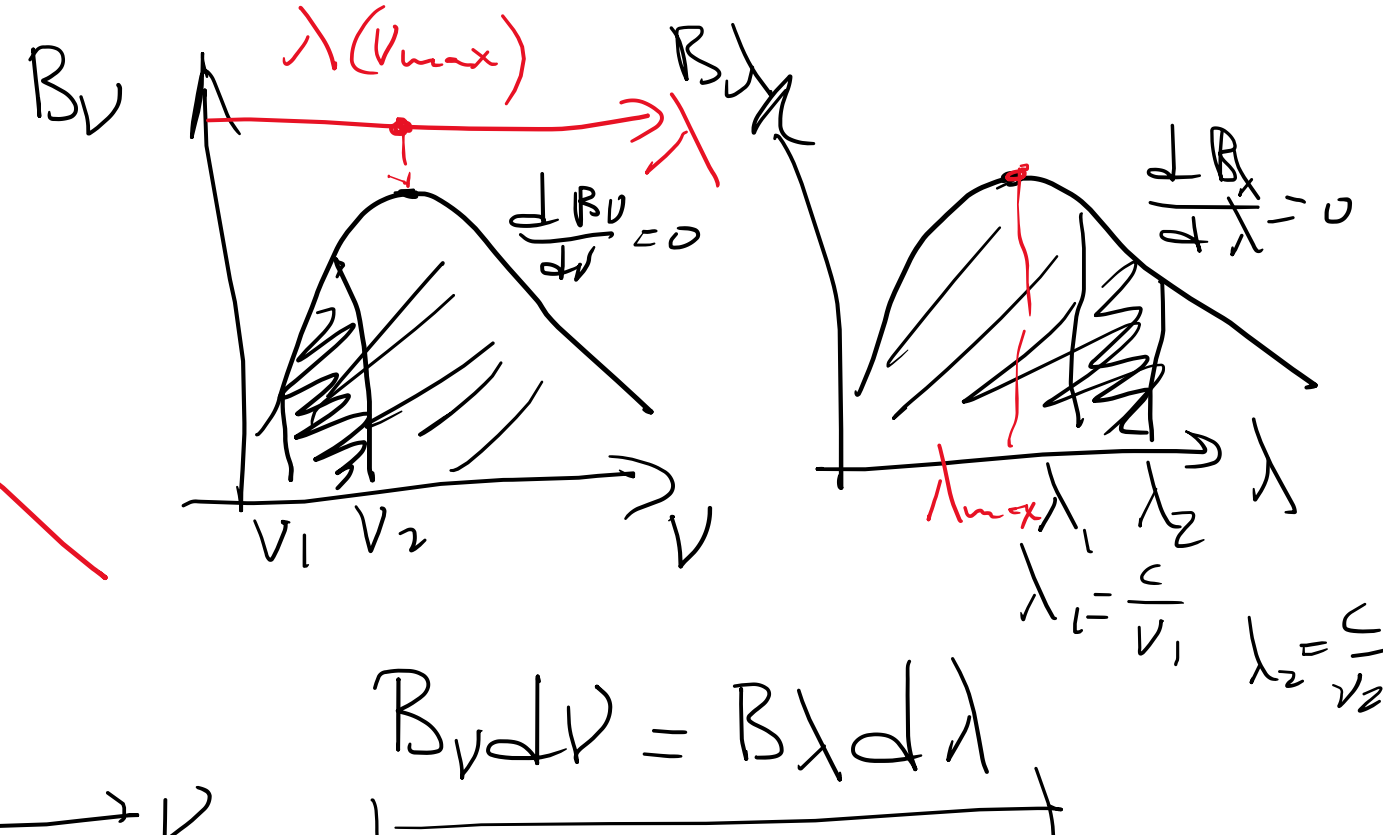
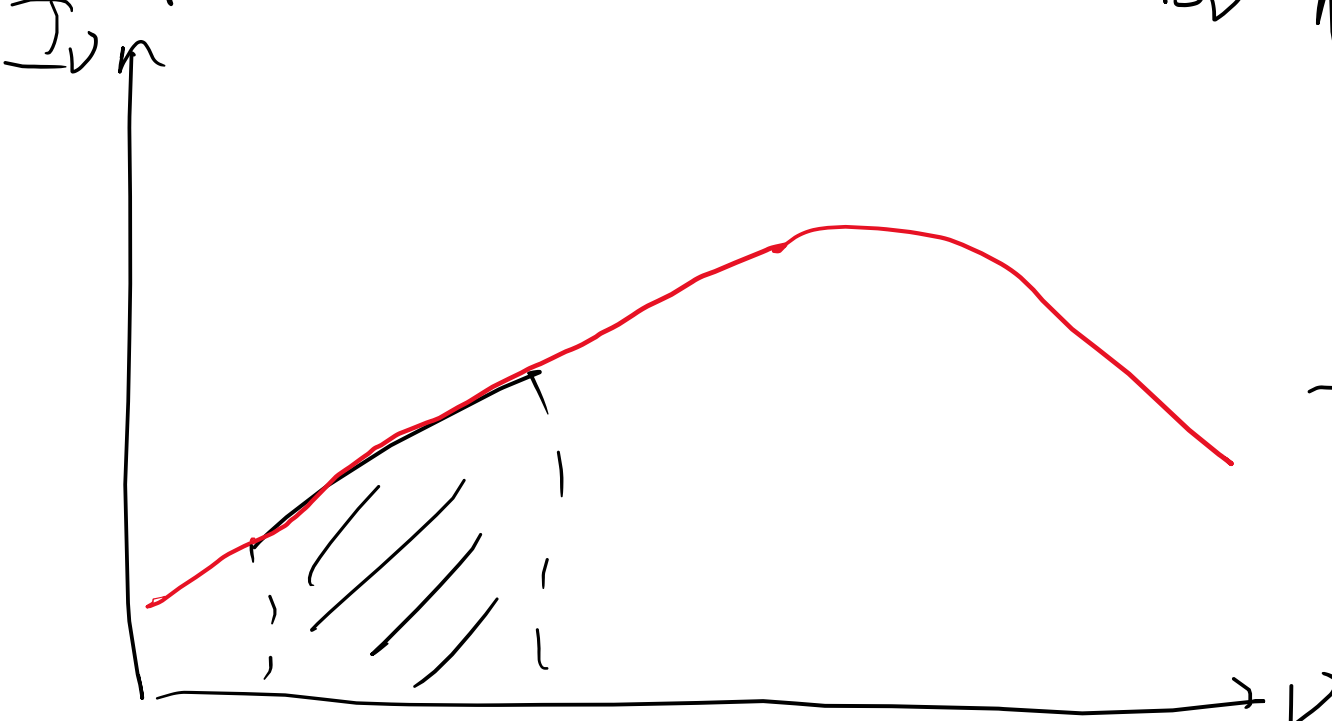
$$S = \pi r^2$$

$$V \approx 10^8 \text{ ly}$$

$$\begin{aligned} 2200 \text{ nm} &= \\ &= 2,2 \cdot 10^3 \cdot 3,1 \cdot 10^{18} \approx 6,7 \cdot 10^{21} \\ &\approx 4 \cdot 10^{43} \text{ cm}^2 \end{aligned}$$



$\rho_{\text{PK-Term}}$



$$\lambda_{\text{max}} \neq \frac{c}{v_{\text{max}}}$$

$$\lambda(v_{\text{max}}) \neq \lambda_{\text{max}}$$

Уловляя термин.

$$\eta_{\text{лет}} = \frac{\text{мощность } (V_1 \div V_2)}{\text{мощность } (V_3 \div V_4)}$$

$$\eta_{\text{лет}}(\nu) = \frac{I_{\nu+\Delta\nu}}{I_{\nu}}$$

$$I_{\nu+\Delta\nu} = I_{\nu} + \Delta\nu \frac{dI_{\nu}}{d\nu}$$

$\Delta\nu \approx \nu$

$$\eta_{\text{лет}} = 1 + \frac{dI_{\nu}}{I_{\nu}} \frac{\nu}{d\nu} = 1 + \frac{d \lg I_{\nu}}{d \lg \nu}$$

Почему:  $B_{\nu} \sim \nu^3 e^{-h\nu/kT}$  (РЛН)

$$\frac{d \lg B_{\nu}}{d \lg \nu} = 3 - \frac{h\nu}{kT}$$

Почему так

$$I_{\nu} \sim \nu^{-\Gamma}$$

$$\frac{d \lg I_{\nu}}{d \lg \nu} \sim -\Gamma$$

$$\nu \sim 3 \cdot 10^8 \text{ Гц}$$

$$\Gamma = 2$$

$$3 - \frac{h\nu}{kT} = -2$$

$$T = \frac{h\nu}{5k} \sim 0,003 \text{ К}$$