



# Майнор. Астрофизика.

Семинар 7. (14.10)

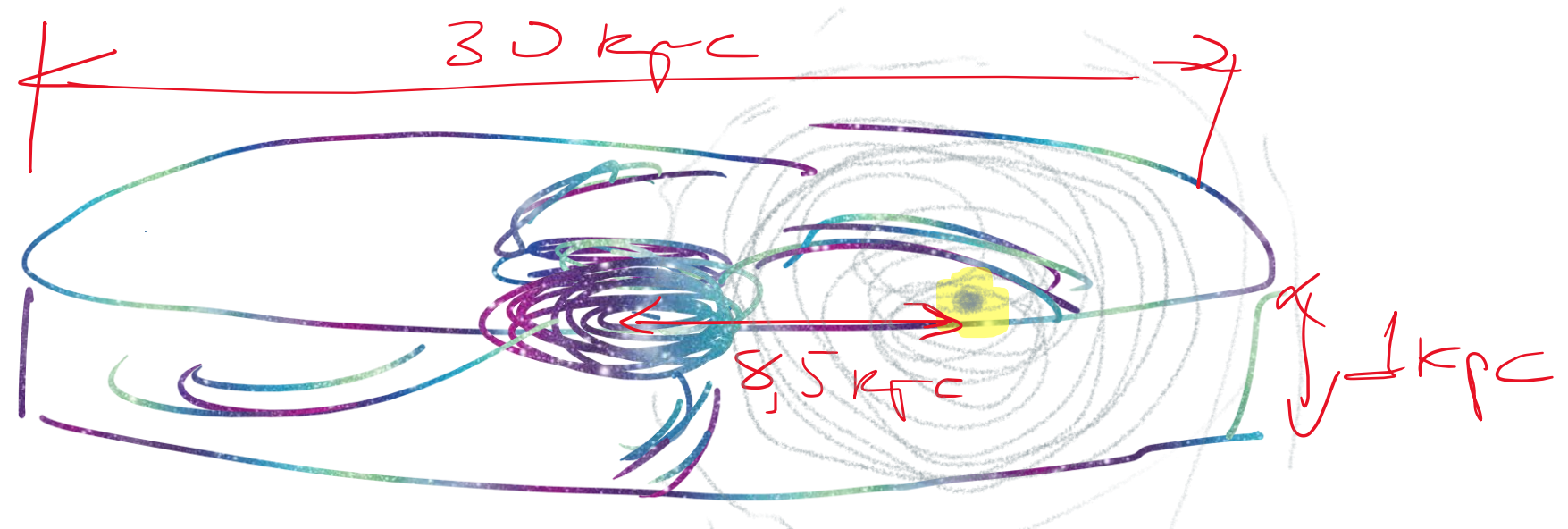
① Παράλλαξη

GAIA 5 kpc 10%  $V = 15^m$

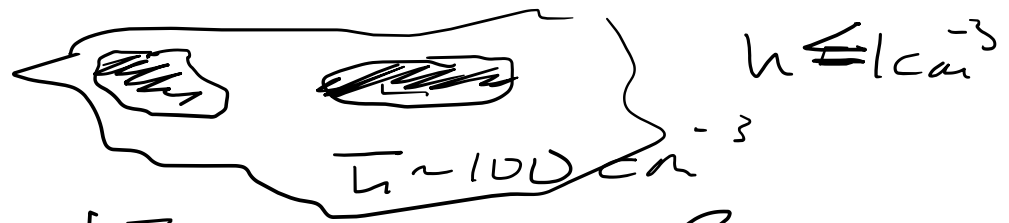
Hipparcos 100 pc 10%  $\rightarrow$  1 kpc

5 kpc:  $\Rightarrow \frac{1}{5}'' = \frac{1}{5000} = 0.0002''$

100 pc  $\Rightarrow \frac{1}{4}'' = \frac{1}{100} = 0.01''$



② Jean's S, T



a)  $E_{\text{TOT}} = E_{\text{kin}} + E_{\text{pot}}$   
 $\begin{matrix} > 0 & < 0 \end{matrix}$

$\begin{cases} E_{\text{kin}} > 0 & \text{ne clez.} \\ E_{\text{TOT}} < 0 & \underline{\underline{\text{clez.}}} \end{cases}$

$2E_{\text{kin}} + E_{\text{pot}} = 0$

$2 \cdot \frac{3}{2} kT \cdot N = \frac{3}{5} \frac{GM^2}{R}$

$N = \frac{M}{m_0}$

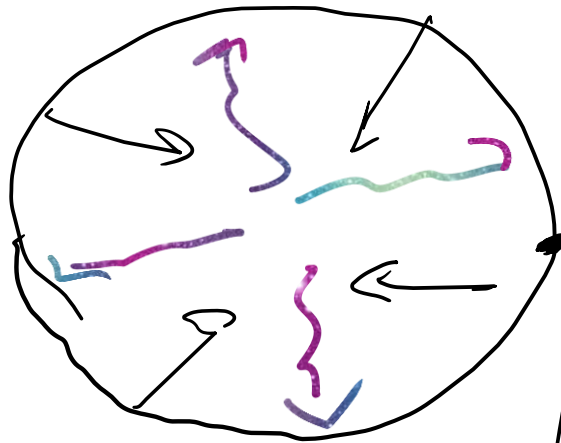
$M = \frac{4}{3} \pi R^3 \rho$

$\frac{kT}{m_0} = \frac{1}{5} \frac{GM}{R} \Rightarrow \frac{1}{5} \frac{G}{R} \frac{4}{3} \pi R^3 \rho = \frac{kT}{m_0}$

$R_J = \left( \frac{15 kT}{4 \pi G m_0 \rho} \right)^{1/2}$

$M_J = \frac{5 kT}{6 m_0} R_J = 3 M_{\odot} \quad \frac{+}{10^4} \quad \frac{R_J}{10^{15} \text{ cm}}$

8)



$$t_{ff} = \frac{R}{v_{ff}(R)} \sim \frac{1}{\sqrt{\epsilon \rho}}$$

$$t_s = R/c_s$$

$$PV = \frac{M}{\mu} RT \Rightarrow$$

$$c_s^2 = \frac{\partial P}{\partial \rho} \sim \frac{P}{\rho}$$

$t_{ff} = \frac{1}{v_{ff}}$

$$c_s = \sqrt{\frac{R}{\mu} T}$$

$$R = k \cdot N_A$$

$$N_A = \frac{\mu}{m_0}$$

$$P = \frac{M}{V} \frac{RT}{\mu} = \rho \frac{RT}{\mu}$$

$$c_s = \sqrt{kT/m_0}$$

$$\frac{R_0}{c_s} = \frac{1}{\sqrt{\epsilon \rho}}$$

$$R_0 = \frac{c_s}{\sqrt{\epsilon \rho}} = \sqrt{\frac{kT}{m_0 \epsilon \rho}} = \sqrt{\frac{RT}{\mu \epsilon \rho}}$$

$$M_0 = \frac{4}{3} \pi R_0^3 \rho$$

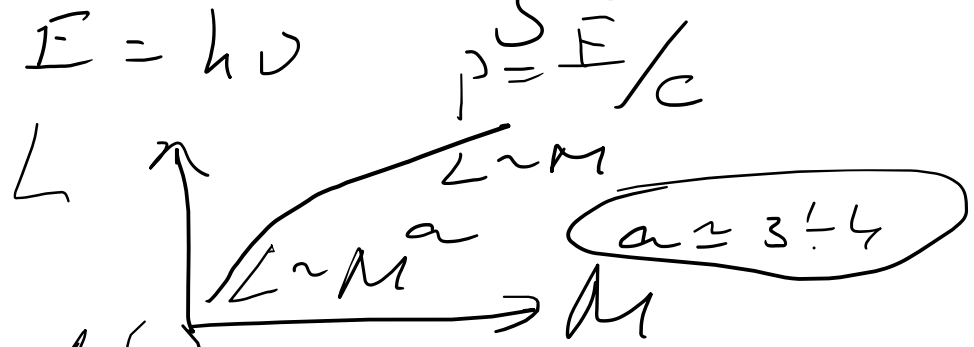


3 M-L

Далее: rayoloe + cleroe

Manomacculore.

$$\frac{dP}{dr} = - \frac{6M(r)g(r)}{r^2}$$



Supra  $E_{tot} = -3PV$

$$P = \frac{-E_{tot}}{3V}$$

$$E_{tot} = - \frac{6a^2}{R}$$



$$P = \frac{6M^2}{4\pi R^4}$$

$$P = nkT$$

$$n = \frac{\rho}{m}$$

$$R = \left( \frac{3}{4\pi \rho} M \right)^{1/3} \sim M^{1/3}, \quad \rho \approx \text{const}$$

$$P = \frac{\rho}{m_0} kT \Rightarrow kT = \frac{6M^2}{4\pi R^4} \frac{m_0}{\rho} = \frac{6Mm_0}{3R}$$

$$L = \frac{4\pi R^2 \sigma T^4}{4\pi R^2} = \frac{4\pi}{3} \left( \frac{3M}{4\pi R} \right)^{2/3} \sigma \left( \frac{6M_{\text{core}}}{3kR} \right)^4$$

$$L \sim M^{2/3} M^4 M^{-4/3} \sim M^{3\frac{1}{3}}$$

$\rho = \text{const}$   
 gelb -  
 2 ziele

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$$L = \frac{N_{\text{year}} \cdot E_{\text{year}}}{\text{year}}$$

$$\tau_{\text{TP}} \sim M/L \sim \frac{M}{M^{3\frac{1}{3}}} \sim M^{-2}$$

⊙  $\tau_{\text{TP}} \sim 10^{10} \text{ a}$

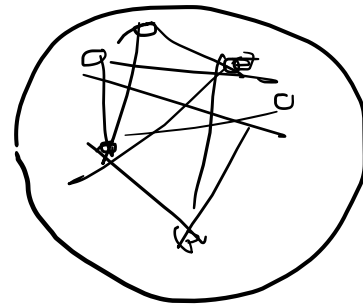
⊕  $\sim 4,568 \text{ Myr}$   
 a)  $\tau_{\text{TP}}$

$2M_{\odot} \tau_{\text{TP}} \sim 2 \text{ Myr}$

$\tau_{\text{TP}} \lesssim 14 \text{ Myr}$

5) M-R

$$h \rightarrow h_c \\ m=1 \quad m=4$$



$$\bar{E}_k \sim \bar{E}_p \text{ of}$$

$$E_k = \frac{3}{2} kT \cdot \omega \sim \frac{M}{m_0} kT$$

$$E_p = \iint \frac{G dm_1 dm_2}{r_{12}} \sim \frac{GM^2}{R}$$

$$\frac{M}{m_0} kT \sim \frac{GM^2}{R}$$

$$R \sim M/T$$

$$T = 1.5 \cdot 10^9 \text{ K} \left( \frac{M}{M_0} \right)^{1/3} \sim M^{1/3}$$

$$\underline{R \sim M^{2/3}}$$



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Пыльсаучуу.

$c_s(r)$

$$c_s^2 = \frac{\partial P}{\partial \rho} \Rightarrow c_s \sim \sqrt{P/\rho}$$

$$P \sim \frac{\epsilon M^2}{R^4} \quad \rho \sim M R^{-3} \quad c_s \sim \sqrt{\frac{\epsilon M}{R}}$$

$$P_{puls} \sim R/c_s \sim \frac{1}{\sqrt{\epsilon \rho}}$$

$$G \cdot P_{puls} = 2\bar{u} \sqrt{\frac{l}{g}} = \dots \quad l \sim R \quad g \sim \frac{\epsilon M}{R^2} \quad \therefore$$

$$= 2\bar{u} \sqrt{\frac{R^3}{\epsilon M}} \sim \frac{1}{\sqrt{\epsilon \rho}}$$



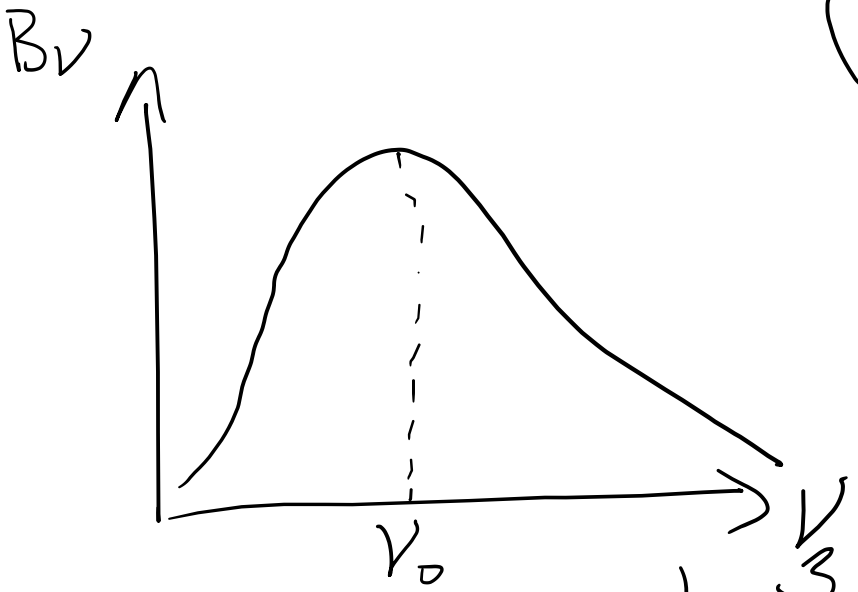
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# Blackbody

$\lambda, \nu, c$

$$c = \lambda \cdot \nu$$

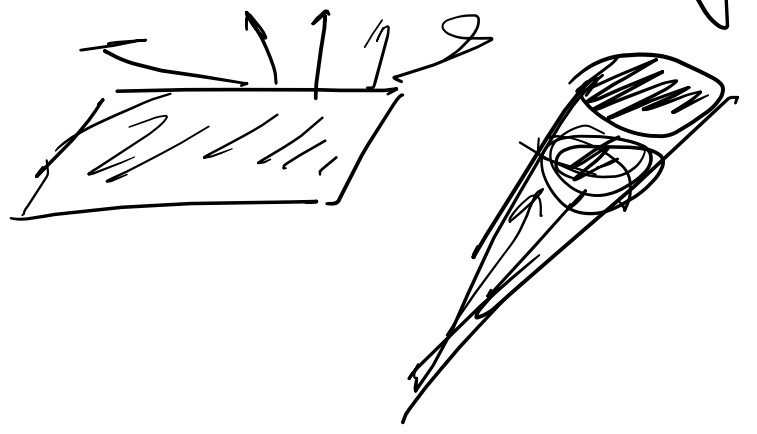
$$E = h\nu$$

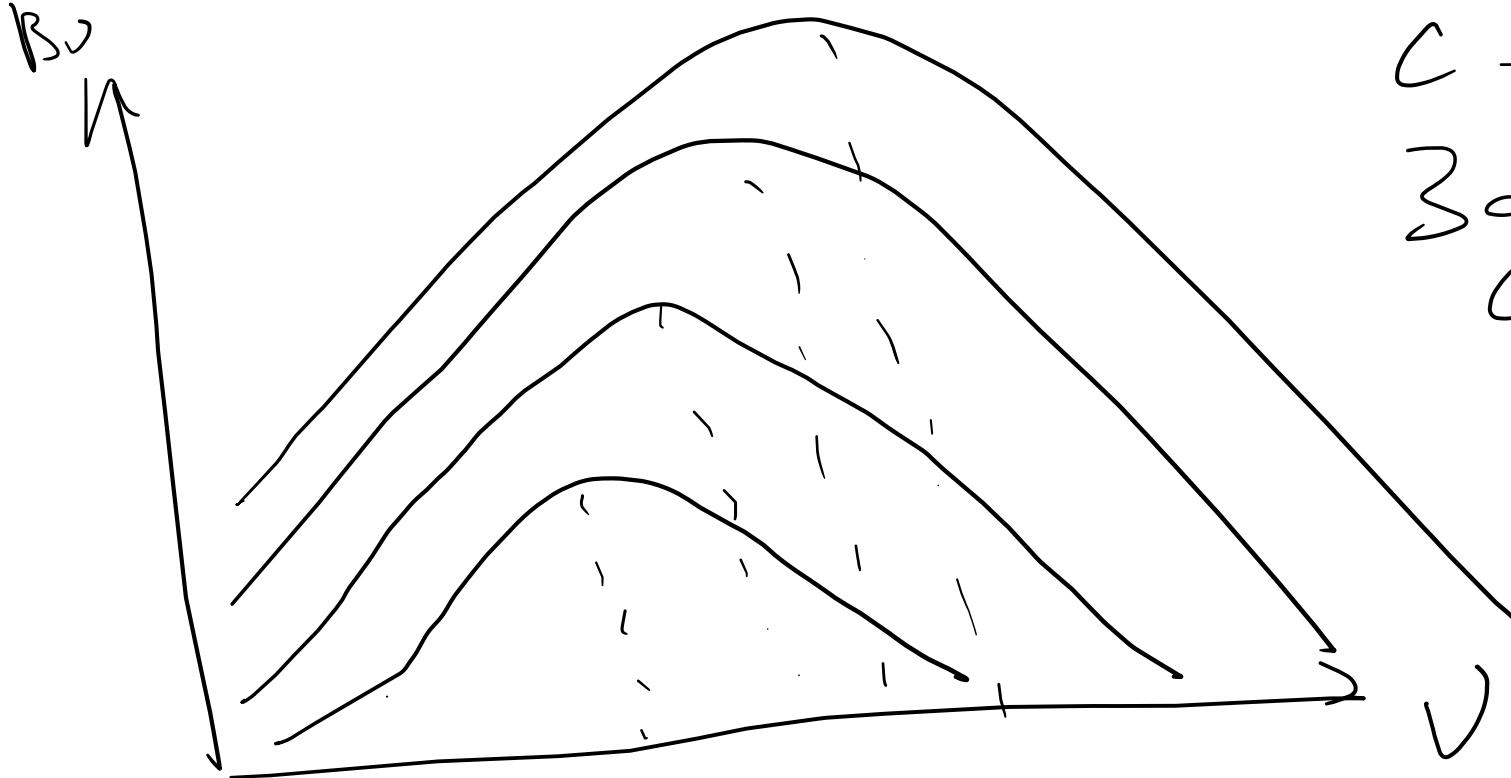


$$[B_\nu] = \frac{\text{дж/с}}{\text{см}^2 \cdot \text{Гц} \cdot \text{стерадиан}}$$

$B_\lambda$

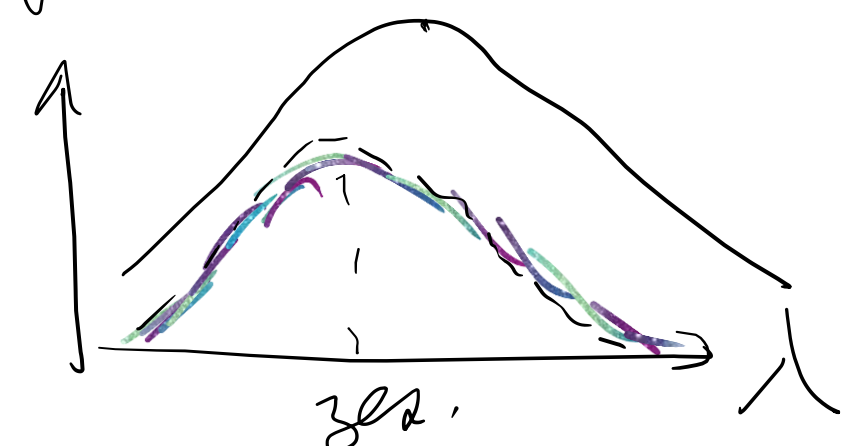
$$B_\nu = \frac{2h\nu^3}{c^2} \frac{1}{e^{h\nu/kT} - 1}$$





$c = \lambda \nu$   
 Закон смещ. Вундта  
 $\lambda_{peak} \sim \frac{1}{T}$

$$\nu_{peak} = T \cdot 5,875 \cdot 10^{11} \frac{T_4}{K}$$



$$E_{peak} = 3kT$$

$10^4 K \rightarrow \sim 1 \text{ эВ}$

Величина энергии сеперка.

$$E_{\text{peak}} \sim 200 \text{ eV}$$

$$T \sim \frac{1}{3(k)} E_{\text{peak}} \sim 70 \text{ eV} \rightarrow \underline{\underline{7 \cdot 10^5 \text{ K}}}$$

$$B(\nu, T) = \frac{2h\nu^3}{c^2} \frac{1}{e^{h\nu/kT} - 1}$$



1) Равенство Димурк  $\nu \rightarrow 0$

$$e^{h\nu/kT}$$

$$\lim_{\nu \rightarrow 0}$$

$$= 1 + \frac{h\nu}{kT} + \dots$$

$$e^x = 1 + \frac{x}{1!} + \frac{x^2}{2!} + \dots$$

$$e^{\frac{h\nu}{kT}} - 1 \approx \frac{h\nu}{kT}$$

$$B(\nu, T) = \frac{2h\nu^3}{c^2} \frac{kT}{h\nu} = \frac{2kT\nu^2}{c^2}$$

Bunoldichte  $\times$  Licht

$$\nu \rightarrow \infty$$

$$e^{h\nu/kT} \gg 1$$

$$B_\nu(\nu, T) = \frac{2h\nu^3}{c^2} \cdot e^{-h\nu/kT}$$

