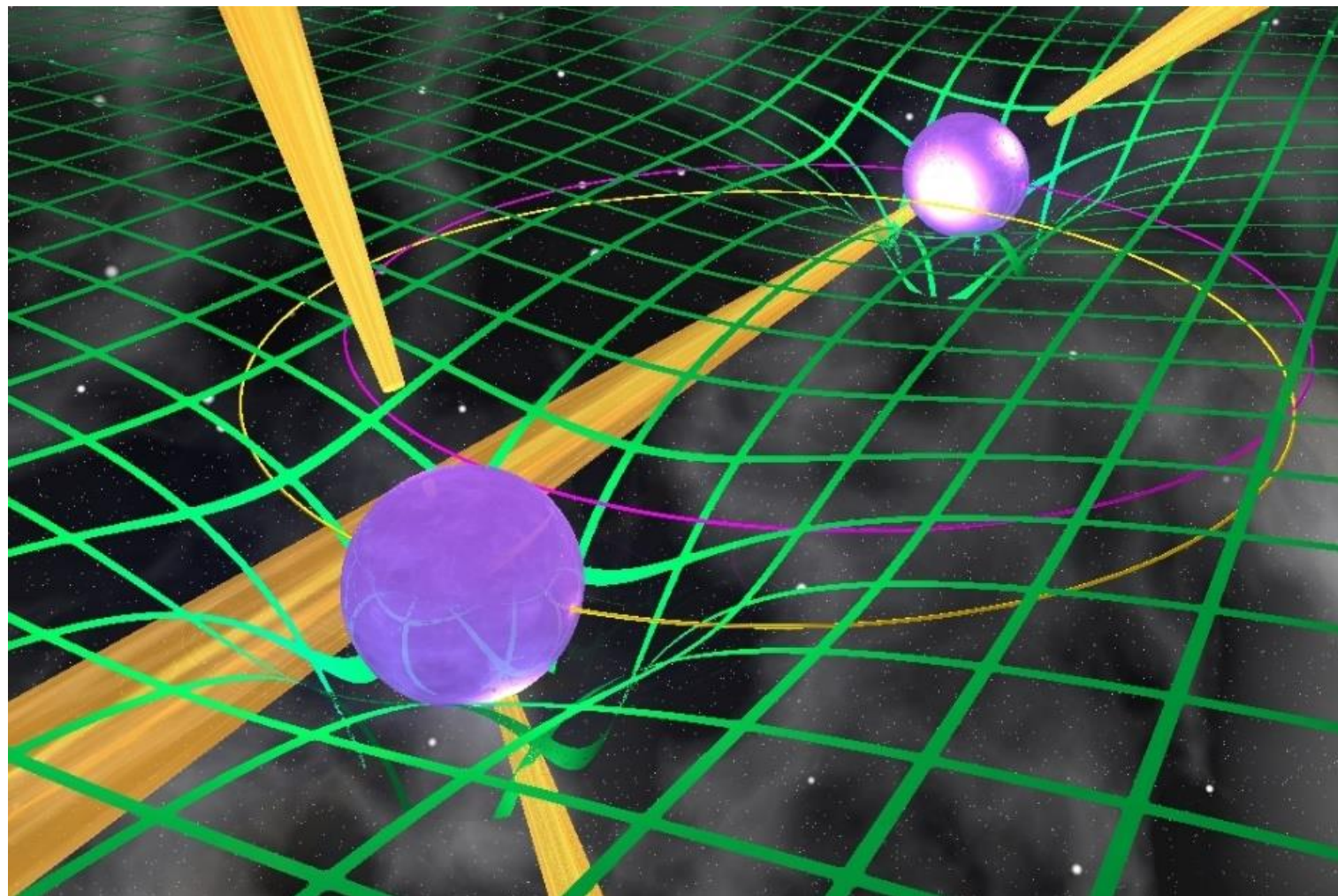
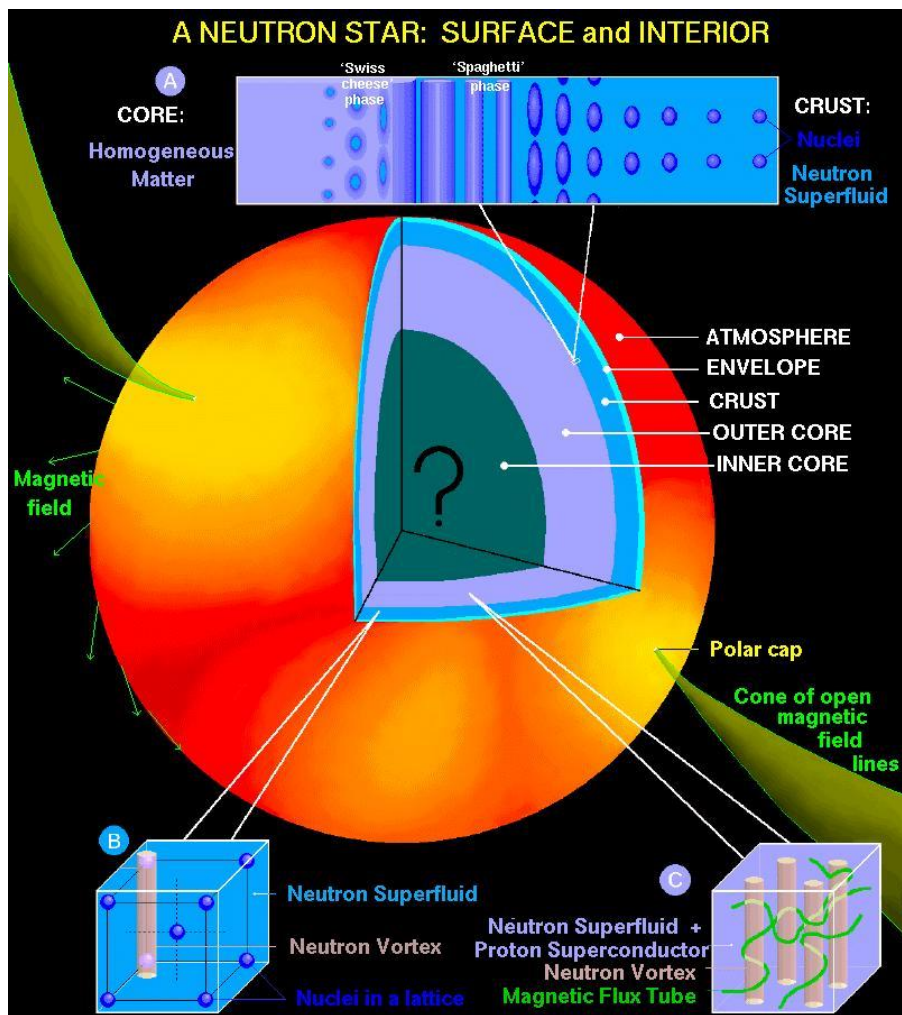




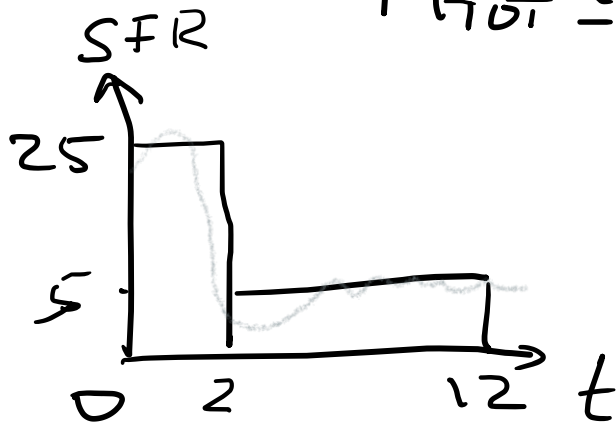
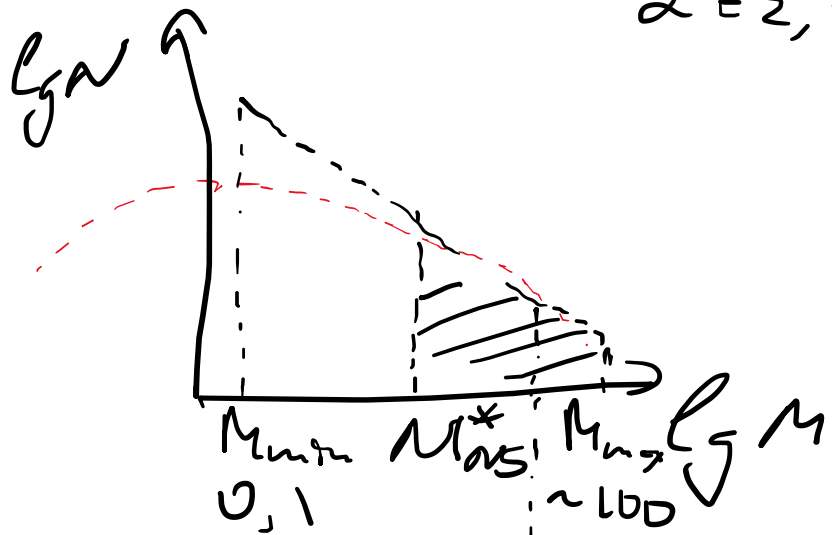
Современная феноменология
нейтронных звезд и черных дыр.

Семинар 1
07.02.2022

①

$$\frac{dN}{dM} = A \cdot M^{-\alpha}$$

$$\alpha = 2,35$$



$$M_{TOT} = 25 \cdot 2 \cdot 10^9 + 5 \cdot 10 \cdot 10 = 10^{11} M_{\odot}$$

$$\langle M \rangle = \frac{M_{NS}}{N}$$

$$= \frac{\int_{0.1}^{100} A \cdot M^{-2.35} M \cdot dM}{\int_{0.1}^{100} A \cdot M^{-2.35} dM} \approx 0.4 M_{\odot}$$

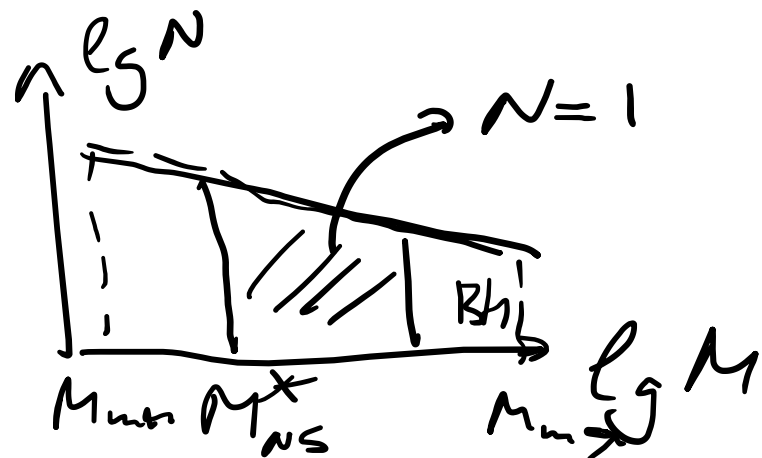
$$\frac{N_{NS}}{N_{TOT}}$$

=

$$\frac{\int_{0.1}^{100} A \cdot M^{-2.35} dM}{\int_{0.1}^{100} A \cdot M^{-2.35} dM}$$

$$\approx \left(\frac{0.1}{10} \right)^{1.35} \approx 10^{-2.7}$$

$$N_{NS} = N_{TOT} \cdot 10^{-2.7} = \frac{M_{TOT}}{\langle M \rangle} = \frac{10^{11}}{0.4} 10^{-2.7} = \frac{2.5}{10^{0.7}} \cdot 10^9$$



$$\int_{10}^{100} A \cdot M^{-2.35} dM = 1 = \tilde{N}_{\text{tot}} \left(\frac{0.1}{10} \right)^{1.35}$$

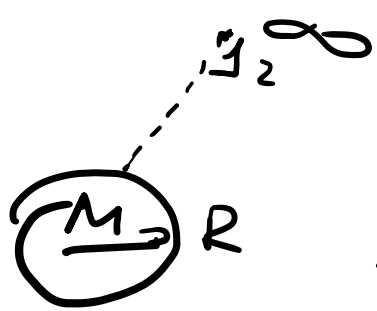
$$\tilde{N}_{\text{tot}} = \frac{\tilde{N}_{\text{tot}}}{\langle M \rangle} = \frac{\overset{= 5 M_{\odot}}{\text{SFR} \cdot T}}{\langle M \rangle} ?$$

$$1 = \frac{5 \frac{M_{\odot}}{208} T}{0.4} \left(\frac{0.1}{10} \right)^{1.35}$$

$$T \approx 40 \text{ ns}$$

2) Излучение NS

a) Акк.



$$M = 1.5 M_{\odot} \quad R = 10 \text{ km}$$

$$E = -\frac{1}{2} \frac{GM}{R} \Rightarrow \frac{1}{2} \cdot \frac{10^{-7}}{1.5} \frac{1.5 \cdot 2 \cdot 10^{33}}{10^4} \approx 10^{20} \text{ эрг}$$

$$L_{\text{Edd}} = 10^{38} \frac{\text{эрг}}{\text{с}} \frac{M}{M_{\odot}}$$

$$L_2 \rightarrow 10^{20} \text{ эрг}$$

$$\sim 10\% \text{ } mc^2$$

$$\dot{M}_{\text{Edd}} \approx 10^{18} \frac{\text{г}}{\text{с}} \approx 10^{-8} \frac{M_{\odot}}{\text{год}}$$

б) Терм. излуч.

$$L = 4\pi R^2 \sigma T^4$$

$$T \sim 10^6 \text{ K}$$

$$L \approx 4\pi \cdot 10^{12} \cdot 5.67 \cdot 10^{-5} \cdot 10^{24} \approx 10^{33} \frac{\text{эрг}}{\text{с}}$$

$$10^6 \text{ K} \rightarrow 100 \text{ эВ}$$

$$\lambda = \frac{0.29(\text{см})}{T} \approx 3 \text{ нм} \rightarrow \text{рентг.}$$

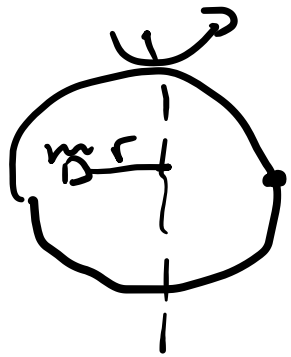
g) Brauerne

$$E_{\text{kin}} = \frac{I \omega^2}{2}$$

$$\omega = \frac{2\pi}{P}$$

$$I \sim MR^2$$

$$I = \underbrace{k}_{\approx 1} MR^2 = \underbrace{k \cdot 3 \cdot 10^3 \cdot 10^{12}}_{\approx 1} \approx 10^{45}$$



$$v_l = v_1 \quad v_l = \frac{2\pi R}{P} = \omega R$$

$$v_1 = \sqrt{\frac{GM}{R}}$$

$$\frac{2\pi R}{P_{\text{min}}} = \sqrt{\frac{GM}{R}} = \sqrt{4} \cdot 10^{-4}$$

$$P_{\text{min}} = 2\pi \frac{R^{3/2}}{\sqrt{GM}} \approx 1 \text{ msec}$$

$$E_{\text{kin}}^{\text{max}} = \frac{10^{45}}{2} \frac{4 \cdot 10^{-2}}{10^{-4}} \approx 2 \cdot 10^{52} \text{ erg} \quad \parallel \quad \omega = \frac{E_{\text{kin}}^{\text{max}}}{E_0} \approx$$

$$E_{\text{kin}} \sim \omega^4 B^2$$

$$\approx \frac{2}{3 \dots} 10^{29} \text{ dex}$$

$$\approx \underline{\underline{2 \cdot 10^{11} \text{ dex}}}$$

2) Max. wave.

$$\mathcal{E} = \frac{B^2}{8\pi}$$

$$E_{\text{mag}} = \mathcal{E} \cdot V = \frac{B^2}{8\pi} \frac{4}{3}\pi R^3 = \frac{B^2 R^3}{6} =$$

$$\omega = \frac{E_{\text{mag}}}{L_0} = \frac{10^{46} B_{14}^2}{6 \cdot 4 \cdot 10^{33} \cdot 3 \cdot 10^7} = \frac{1}{6} \cdot 10^{46} \text{ s}^{-1} B_{14}^2$$

$$\mathcal{E} = \rho_B \cdot c^2$$

$$\rho_B = \frac{B^2}{8\pi c^2} = \frac{10^{30} B_{15}^2}{8\pi 10^{21}} \approx \underline{\underline{4 \cdot 10^7 \text{ g/cm}^3}}$$

B_{max} :

$$\mathcal{E} \cdot V = E_p$$

$$\frac{B_{\text{max}}^2 R^3}{6} = \frac{6 M^2}{R}$$

$$B_{\text{max}} = \left[6 \frac{6 M^2}{R^4} \right]^{1/2} = \left[6 \cdot \frac{10^{-7} \cdot 9 \cdot 10^{46}}{1.5 \cdot 10^{24}} \right]^{1/2} = 6 \cdot 3 \cdot 10^{17} \text{ G} \\ \approx 2 \cdot 10^{18} \text{ G}$$